28. HYDROCARBON GASES IN CANNED CORE SAMPLES FROM LEG 28 SITES 271, 272, AND 273, ROSS SEA

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INTRODUCTION

Samples from the Ross Sea cores are among the richest yet analyzed for hydrocarbon gases; and they contain small but significant amounts of ethane and heavier gaseous hydrocarbons. The gas was anticipated because of shipboard observations of its presence. This observation also was responsible for the greater sampling density, particularly at Site 272.

As in previous cases the samples were collected soon after recovery and sealed in ordinary size 2½ tin cans. Upon receipt at the laboratory, 21 cans were chosen for these analyses. A few of the cans had already begun leaking—evidenced by salt incrustation and corrosion at one or more points along the edges of the lid. Some of these leaking cans were found to contain virtually no hydrocarbon gas, so each can was carefully inspected, and the few bad ones were set aside.

The selected cans were punctured through a rubber septum clamped tightly to their sides. A measured volume of each can's air space was removed with a hypodermic syringe and introduced into the inlet system of a Perkin-Elmer Model 154 B gas chromatograph with a hydrogen-flame detector. The column is a 2-meter P.E. column B (di-2-ethylhexyl sebacate); temperature, 46°C; helium pressure, 8 psig; hydrogen pressure, 8 psig. From the volume of gas and sediment in the cans, and the gas chromatographic results, the gas contents were calculated as parts per million of gas by volume (i.e., mls of gas at STP per 10⁶ mls of sediment as received). After gas-space samples were run in duplicate, the cans were opened and a sample of sediment was removed and agitated with water in a Waring blender to remove any gas remaining in the sticky mud. The gravities were so small (a few ppm of methane at most), that after the first few samples only the can's free (gas-space) gas was analyzed subsequently.

RESULTS AND DISCUSSION

Results of organic carbon and gas determinations are listed in Table 1. They are also compared, diagrammatically, with all samples run previously, by the histograms in Figures 1 and 2. Overall, muds from Holes 272 and 273A have more gas than any other samples run to date; that is, there is more residual gas remaining in the cans with them, on the average, than with samples from any other suite. They are more consistently high in value than Site 186 (McIver, 1973) that had one value of 216,700 ppm of methane. Site 229 (McIver, 1974) had one sample with 165,000 ppm methane, but the second was 84,500 ppm. Only two

Sample (Interval in cm)	Organic Carbon (% by wt)	Total Hydro- carbon Gas (ppm by vol.)	Ethane and Heavier HC Gas (ppm by vol.)	Ratio, Total HC Gas to Ethane +
271-24-2, 50-70	0.25	3,700	111	33
272-7-2, 140-150	0.53	15,900	71	224
272-8-1, 0-20	0.32	26,900	113	238
272-10-4, 0-10	0.36	68,000	138	492
272-13-3, 0-10	0.38	38,900	131	297
272-14-4, 0-6	0.46	115,000	76	1513
272-18-1, 143-150	0.43	50,400	62	813
272-21-5, 144-150	0.44	29,500	tr	>1400
272-23-3, 140-150	0.46	170,300	177	962
272-27-2, 137-144	0.41	145,100	274	529
272-29-2, 140-150	0.39	75,500	100	686
272-31-1, 0-5	0.39	179,000	58	3086
272-33-2, 0-8	0.41	54,000	tr	>2700
272-39-1, 0-5	0.76	27,700	38	728
273A-11-4, 140-150	0.47	52,000	90	578
273A-13-4, 140-150	0.51	73,000	716	102
273A-14-2, 6-16	0.46	132,000	112	1179
273A-16-2, 0-7	0.47	147,000	876	168
273A-17-2, 140-150	0.45	111,200	88	1263
273A-22-2, 143-150	0.05	136,000	636	214
273A-25-2, 144-150	0.38	60,000	77	787

 TABLE 1

 Analytical Results – Leg 28 Samples

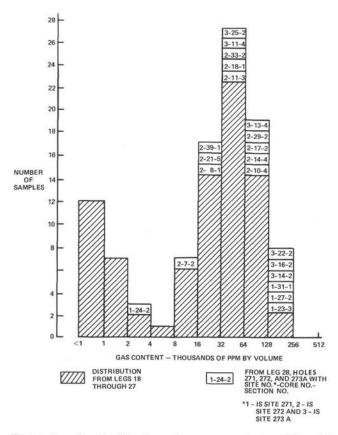


Figure 1. Residual hydrocarbon gas contents of Leg 28 samples compared with preceding legs.

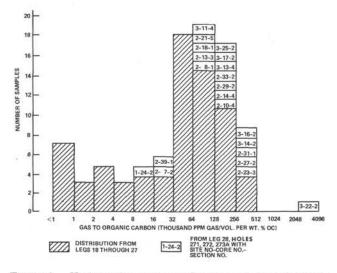


Figure 2. Hydrocarbon gas as a function of organic carbon contents – Leg 28 compared with previous results.

samples were sent from this location so we cannot say anything about uniformity of methane contents down the hole.

Perhaps even more important is the small but significant quantity of ethane and higher homologs appearing in some of these gases, particularly those from Hole 273A. In all samples from previous suites of canned cores, the amount of ethane-plus was generally very low. The background of a procedure is of the order of 20 ppm and seldom until this suite did more appear. (A few notable exceptions are tabulated in Table 2.) Ethane is generally attributed to thermal diagenesis, or maturation, of organic matter at depth, and in DSDP cores may be considered as evidence that somewhere in the vicinity, more deeply buried rocks are already experiencing the early stages of thermal cracking that leads to the generation of light hydrocarbons (McIver, 1972).

Generally in the DSDP, the ratio of hydrocarbon gas to its heavier homologs in the canned gases was greater than 1000 and often at least 3000 or 4000. If this is "background", then lower ratios may be evidence of local organic diagenesis. The erratic values of this ratio, from 200 to 3000 at Site 272, suggests irregular dilution of the methane gas from migrating ethane. The generally lower values (100-1200) at Site 273A suggest more abundant generation and migration. And the value, 33, at Site 271 is evidence of very extensive ethane generation and migration. These ratios are generally lower, but ranked in the same order as virtually the same ratio, methane to C2H6, determined onboard Glomar Challenger on samples taken from gas "gaps" in the core, observed through the core liner. The values thus obtained were 75 and 88 for Site 271, 2000 to 3000 at Site 272, and 500 to 700 at Site 273A. The differences between results from the two analog sites may be attributed to different methods including the timing of collecting the samples. Those provided us had probably bled off their gas for longer periods and there may have been a more rapid (or preferential) release of methane, particularly if the gases were immobilized as hydrates at these sites.

TABLE 2 Ethane + Contents of Residual, Canned Gases in Excess of 50 ppmv of Total Cuttings

Leg	Site	Core	Section	Ethane + Content (ppmv)	Total HC Gas (ppmv)	Ratio, total HC Gas to Ethane +
18	174A	15	4	122	24,300	199
18	174A	40	5	209	41,400	198
18	176	(none)				>1000
18	180	(none)				>1000
19	185	(none)				>1000
19	186	(none)				
19	189	(none)				
19	191	(none)				
21	204	(none)				>1000
21	210	(none)				
23	222	(none)				>1000
23	229	(none)				
23	229A	(none)				
24	231	13	4	78	42,000	538
24	231	13	4 5	95.	67,000	705
24	231	16	4/5	246	62,600	254
24	232	(none)	0.02552.0		1.1997 - 1897 - 199 - 1997 -	
27	262	(none)				>3000
28	(see Ta	ble 1)				

That these samples from Holes 271, 272, and 273A are very gassy is also evidenced by their gas contents as a function of their organic carbon contents. As Figure 2 shows, these samples have more gas per percent organic carbon, on the average, than the samples from the other 15 sites sampled to date. The samples have only moderate amounts of organic matter; but whether the nature of the organic matter is different; i.e., more susceptible to bacterial attack than the commonly found type; or whether this abundance simply reflects more intensive migration, cannot be determined. It appears that this would be a very interesting area for more extensive organic geochemical studies.

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