24. MELTING RELATIONS AND VISCOSITY OF AN ABYSSAL OLIVINE THOLEIITE

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ABSTRACT

Melting experiments at 1 atmosphere on an olivine tholeiite, Sample 395A-8-1, 50 cm, suggest that the magma erupted at or near 1215 °C. On the basis of the experiments at high pressures up to 12 kbar, it appears that olivine tholeiite magma could be equilibrated with plagioclase lherzolite at a depth of about 25 km or be formed by fractional crystallization from more magnesian magma at depths shallower than 25 km. It also appears that the process of fractional crystallization is most likely at a depth of 25 km if pyroxenes take an important part in such a process.

Viscosity and density measurements of melt of the sample at high pressure suggest that owing to low viscosity, olivine and pyroxene sink rapidly in the melt, whereas plagioclase floats. The phenocrysts of porphyritic basalt, therefore, do not always represent the first liquidus phase. Some plagioclase-phyric tholeiite may be derived from olivine tholeiite magma by floating of plagioclase phenocrysts at high pressures.

INTRODUCTION

Petrography and geochemistry of oceanic basalts have been studies by many investigators (e.g., Bryan et al., 1976). Experimental data on such rocks are relatively scarce, however. It is important to know the melting relations of oceanic basalt at 1 atmosphere and at high pressures, in order to understand the genesis of oceanic basalt magmas. Viscosity and density data on oceanic basalt magma at high pressures are also necessary to understand processes of fractionation at high pressures. For these reasons, melting experiments and density and viscosity measurements have been conducted on melt of a sample collected near the Mid-Atlantic Ridge during Leg 45.

Sample 395A-8-1, 50 cm, was recovered at a depth of about 140 meters beneath the ocean floor, and belongs to type A_2 aphyric tholeiite. This sample is glassy olivine tholeiite, probably from the chilled margin of a pillow basalt, containing very small amounts of olivine microphenocryst (0.3 mm across, Fo₈₅) and platy skeletal crystals of olivine (0.1 mm long, Fo₈₃) set in the glassy matrix. The modal amount of olivine microphenocrysts is less than 1 per cent. The microphenocryst also shows skeletal habit, indicating that the microphenocrysts grew rapidly probably just before or at the time of eruption of magma. It may thus be concluded that this sample represents the liquid composition.

The chemical composition and CIPW norm of the sample are listed in Table 1. This sample is one of the most magnesian rocks (Mg/Mg+Fe^{*}=0.57) among the aphyric olivine tholeiite recovered on Leg 45, and also has high Ni content (178 ppm), low K₂O content (0.13 wt%), and low ignition loss (Bougault et al., this volume). It therefore appears to be the least fractionated and least altered olivine tholeiite recovered during Leg 45.

MELTING EXPERIMENTS

The 1-atmosphere melting experiments were performed at $pO_2 = 10^{-8}$ bar, where pO_2 was controlled by changing the mixing ratio of CO_2 and H_2 gases. Powdered samples were pressed into pellets and a fragment of the pellet tied with fine Pt wire was fused at 1300°C and $pO_2 = 10^{-8}$ bar for 15 minutes to make a glass sphere. These glass spheres, held on fine Pt wire, were used through runs in order to minimize iron loss. Runs were quenched by dropping the material into a metal cup cooled with ice from outside.

The high-pressure melting experiments were carried out in graphite capsules in a solid-media apparatus (Boyd and England, 1960) using the piston-out method. The powdered samples were dried at 110°C for several days and heated in air in graphite capsules at about 1000°C for a minute just before the experiment to eliminate water. The oxygen fugacity of each run was within the stability field of wüstite (Thompson and Kushiro, 1972).

The results of melting experiments are listed in Tables 2A and 2B, and plotted in Figure 1. Olivine is the liquidus phase at pressures below 8 kbar. Above this pressure, Ca-rich clinopyroxene appears on the liqui-

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Composition		CIPW norm		
SiO2	50.0	Or	0.78	
TiO ₂	1.64	Ab	22.75	
Al ₂ O ₃	15.0	An	28.23	
FeO ^a	11.2	Di	19.83	
MnO	0.22	Hy	14.53	
MgO	8.40	01	10.79	
CaO	10.6	11	3.09	
Na ₂ O	2.70			
к ₂ 0	0.13			
Total	99.9			

TABLE 1 Chemical Composition and

^aTotal iron as FeO.

TABLE 2A Results of One-Atmosphere Experiments on an Olivine Tholeiite, Sample 395A-9-1, 50 cm (pO₂ = 10⁻⁸ bar)

Temperature (°C)	Time (hr)	Products		
1220	19	Gl		
1210	20	Ol + Gl		
1200	20	Ol + Gl		
1190	27	Ol + Sp + Gl		
1180	42	Ol + Pl + Sp + Gl		
1170	43	Ol + Pl + Sp + Gl		
1160	42	Ol + Pl + Sp + Gl		
1150	45	Ol + Pl + Cpx + Gl		
1120	64	Ol + Pl + Cpx + Opaque minerals + Gl		
1100	118	Fine crystalline aggregates		

Note: Abbreviations: Ol, olivine; Cpx, clinopyroxene, Pl, plagioclase; Sp, spinel; Gl, glass.

dus. Although spinel is the second mineral to crystallize at 1 atm, it was not detected in the high-pressure experiments. At about 8 kbar, olivine, plagioclase, Carich clinopyroxene, and orthopyroxene crystallized within 10°C below the liquidus, indicating that the liquid lies close to the four-phase boundary. This may also mean that orthopyroxene is on the liquidus at or near this pressure.

DENSITY AND VISCOSITY MEASUREMENTS **AT HIGH PRESSURES**

The densities of melt were measured by a fallingsphere method, using spheres with different densities. To reduce the thermal gradients within the graphite capsule, specially designed furnaces (Kushiro, 1976)

TABLE 2B Results of High-Pressure Melting Experiments on an Olivine Tholeiite, Sample 395A-8-1, 50 cm

Pressure (kbar)	Temperature (°C)	Time (min)	Products		
5	1215	50	Ol + Gl		
7.5	1230	90	Gl		
7.5	1220	90	Ol + Pl + Gl		
7.5	1210	95	Ol + Pl + Cpx + Opx(?) + Gl		
7.5	1175	90	Ol + Pl + Cpx + Gl		
8.2	1225	40	Ol + Cpx + Gl		
8.2	1215	80	Ol + Cpx + Opx + Pl + Gl		
8.2	1205	95	Ol + Cpx + Opx + Pl + Gl		
9	1215	90	Cpx + Opx + Pl + Gl		
10	1270	80	Gl		
10	1250	90	Cpx + Gl		
10	1230	70	Cpx + Pl + Gl		
12	1300	20	Gl		
12	1260	90	Cpx + Gl		

Note: Abbreviations: Ol, olivine; Cpx, clinopyroxene; Opx, orthopyroxene, Pl, plagioclase; Gl, glass.



Figure 1. Phase relations of an olivine tholeiite, Sample 395A-8-1, 50 cm.

were used. Spheres of olivine (8 kbar measurement) or chromian diposide (12 kbar measurement) were placed at the top, and spheres of boron nitride and plagioclase (An₆₅) were placed at the bottom of the charge when it was loaded. The distances the spheres dropped (olivine or chromian diopside) or floated (boron nitride and plagioclase) were measured with a microscale after quenching of each run (Table 3). There was no evidence of reaction between melt and spheres.

The moving velocity, V, of a sphere with radius of rin the melt with viscosity η is

$$V = \frac{2r^2 \cdot \Delta \rho \cdot g}{9\eta} = \frac{\varrho}{\Delta t} \tag{1}$$

TABLE 3 Results of Runs for Density Measurement

Pressure (kbar)	Temperature (°C)	Time (min)	Diameter (mm)	Sinking distance (mm)	Sphere
8	1250	10	0.70	2.6	Ol
. 8	1250	10	0.80	3.3	Ol
8	1250	10	1.0	-1.4	Pl
8	1250	10	0.60	-2.3	BN
12	1300	10	0.50	2.9	Cpx
12	1300	10	0.80	-2.3	Pl
12	1300	10	0.60	-5.0	BN

Note: Abbreviations: Ol, olivine (F091); Pl, plagioclase (An65); BN, fired born nitride, Cpx, chromian diopside.

where *l* is the distance the sphere moved during time interval Δt , ρ is the density difference between spheres and melt, and *g* is the gravity constant. For the spheres with different densities in the same melt, the following relation is derived from Equation 1:

$$\frac{\ell_1}{r_1^2(\rho_1 - \rho_0)} = \frac{\ell_2}{r_2^2(\rho_2 - \rho_0)} = \frac{2 \cdot g \cdot \Delta t}{9 \eta}$$
(2)

where l_1 and l_2 are the moved distances, r_1 and r_2 are the radii, ρ_1 and ρ_2 are the densities of spheres 1 and 2, respectively, and ρ_0 is the density of melt. As three spheres with different densities were used in each run, the density of the melt was calculated by the leastsquares method, on the basis of equation 2. Although the densities of spheres were 3.33 (olivine and chromian diopside), 2.70 (plagioclase, An₆₅) and 2.29 (boron nitride) at 1 atmosphere and at room temperature, 3.23, 2.68, and 2.25 were used in the calculation to take into account the effects of compressibility and thermal expansion (Birch, 1966; Skinner, 1966; Yagi, unpublished data).

The calculated densities were 2.80 g/cm³(± 0.03) both at 1250°C and 8 kbars and at 1300°C and 12 kbar. This value is larger than that calculated from the chemical composition, according to Bottinga and Weill (1970). The difference is about 0.1 g/cm³ when iron is treated as ferrous, and about 0.2 g/cm³ when iron is treated as ferric.

The viscosity measurements were based on the method described by Kushiro et al. (1976), except that olivine spheres were used at 8 kbar instead of chromian diopside spheres. Run time ranged from 3 minutes to 10 minutes.

The sinking velocities were calculated from results of at least two runs with different run times at the same conditions (Table 4). They were 0.36 mm per minute for spheres of olivine with radius of 0.40 mm and 0.28 mm per minute for spheres of olivine with radius of 0.35 mm at 8 kbar and 1250°C; 0.48 mm per minute for spheres of chromian diopside with radius of 0.30 mm at 12 kbar and 1300°C; and 0.2 mm per minute

TABLE 4 Results of Runs for Viscosity Measurement

Pressure (kbar)	Temperature (°C)	Time (min)	Diameter (mm)	Sinking distance (mm)	Sphere
8	1250	6	0.80	1.9	Ol
8	1250	6	0.71	1.5	Ql
8 ^a	1250	10	0.80	3.3	Ol
8 ^a	1250	10	0.70	2.6	O 1
12	1275	3	0.70	0.5	Cpx
12	1275	6	0.72	1.0	Cpx
12	1275	10	0.70	1.8	Cpx
12	1300	5	0.60	1.7	Cpx
12	1300	9	0.60	3.6	Cpx

Note: Abbreviations: Ol, olivine (Fog1); Cpx, chromian diopside ^aFrom Table 3.

for spheres of chromian diopside with radius of about 0.35 mm at 12 kbar and 1275°C (Figure 2). The viscosities of melt were calculated by Stokes' law, using the density data described above, and corrections for wall effects were based on the Faxen correction (Shaw, 1963). The results are shown in Figure 3 with the 1-atmosphere viscosities calculated according to model of Bottinga and Weill (1972). The relatively high viscosity at 12 kbar and 1275°C probably results from the presence of clinopyroxene crystals because it is below the liquidus temperature. The calculated 1-atmosphere viscosity is clearly higher than the viscosity at high pressures at the same temperature.

DISCUSSION

The 1-atmosphere experiments suggest that olivine begins to crystallize at 1215°C in this magma. The ex-



Figure 2. Time and distance relations for sinking of spheres in melts of an olivine tholeiite, Sample 395A-8-1, 50 cm. Solid circle, olivine sphere; open circle and square, chromian diopside sphere.



Figure 3. Viscosity of an olivine tholeiite melt. Numbers in rectangles indicate measured viscosity (poise). Those in parentheses are calculated according to the method of Bottinga and Weill (1972).

istence of a small amount of olivine microphenocrysts in the sample indicates that this magma erupted at or near $1215 \,^{\circ}\text{C}$.

O'Hara (1968) stated that abyssal tholeiite cannot be a primary magma derived by partial melting of mantle peridotite, because most abyssal tholeiites do not coexist with orthopyroxene at any pressure. As shown above, however, the composition of this olivine tholeiite is close to the olivine-plagioclase-Ca-rich clinopyroxene-orthopyroxene-liquid boundary at about 8 kbar. And it may be potentially a primary magma in the sense of O'Hara (1968). O'Hara (1973) also stated that abyssal tholeiite cannot be a partial melting product of mantle peridotite, because the near-liquidus orthopyroxene is not so magnesian as those in mantle peridotities. Unfortunately, it was impossible to obtain the exact analysis of orthopyroxene without contamination of glass, because of its small grain size. The olivine coexisting with orthopyroxene was Fo₈₆. Because the Fe-Mg partitioning between olivine and orthopyroxene is almost unity and the partition coefficient is insensitive to temperature and pressure, the composition of orthopyroxene should be close to En₈₆. Some of the mantle-derived peridotites contain similar olivine and orthopyroxene (Kuno and Aoki, 1970). It is possible that under anhydrous conditions this olivine tholeiite magma can coexist in equilibrium with plagioclase lherzolite of oceanic upper mantle at a depth of about 25 kilometers.

Melting relations similar to the results given above were obtained on an olivine tholeiite dredged from the Mid-Atlantic Ridge at $45^{\circ}N$ (Kushiro, 1973). In that case, the pressure and temperature conditions for the five-phase assemblage (Ol + Cpx + Opx + P1 + Liq) are nearly the same (7.5 kbar, 1220°C), in spite of the different chemical compositions (espeically in MgO and FeO) of the two starting materials. The melting relations of a magnesian olivine tholeiite from the FAMOUS area (Hodges, 1976; Fujii, unpublished data) are also similar, but the liquidus temperature is much higher because of high Mg/Fe ratios. The olivine tholeiite used in this experiment and that by Kushiro (1973) may be products of a lower degree of melting than that of FAMOUS area (Hodges, 1976). It may be concluded, therefore, that some of the abyssal olivine tholeiite may be equilibrated with the upper mantle material at about 8 kbar, and that the compositional variations may result from the difference in degree of melting or different degrees of depletion in the source materials.

Although it was emphasized above that the olivine tholeiite used here may represent a primary magma formed by direct partial melting of upper mantle materials, it is also possible that this magma is a product of fractional crystallization of an even more magnesian primary magma; but such a fractionation process is limited to depths shallower than 25 km where olivine is the first liquidus phase. Considering the decrease of viscosity with increasing of pressure (Figure 3), however, such a fractionation process is more effective at high pressures than near 1 atmosphere.

The diverse compositional variations of abyssal olivine tholeiite have been explained by a process of separation of olivine and plagioclase from a primary magma (Kay et al., 1970; Hart, 1971). Recently, however, Shibata and Fox (1975), Presnall and O'Donnell (1976), and Clague and Bunch (1976) argued that the diversity of abyssal olivine tholeiite cannot be explained by the separation of olivine and/or plagioclase alone, but requires the additional separation of Ca-rich clinopyroxene. The above experimental results suggest that such a differentiation process is possible at about 8 kbar where olivine, pyroxenes, and plagioclase are near-liquidus phases.

Recently, three different types of abyssal tholeiite have been distinguished on the basis of the primary phase: olivine tholeiite, plagioclase tholeiite, pyroxene tholeiite (Shido et al., 1971; Hekinian and Aumento, 1973; Shido et al., 1974; Bougault and Hekinian, 1974). In those arguments, it is assumed that the phenocrysts or microphenocrysts represent the first liquidus phase and the bulk compositions represent the compositions of liquid. The low viscosity (Figure 3) and the high density (2.80 g/cm³) of melt at high pressures, however, suggest that olivine and pyroxenes sink rapidly, whereas plagioclase floats in the magma. Consequently, the phenocrysts do not always represent the first liquidus phase, and the bulk compositions of porphyritic rock do not always represent the compositions of liquid. These considerations suggest that it is possible to derive two different types of tholeiites, plagioclase-phyric tholeiite and olivine-phyric tholeiite, from an olivine tholeiite parent magma. Some abyssal plagioclase-phyric basalts contain huge glomerocrystic plagioclase phenocrysts (xenocryst) and smaller euhedral olivine (e.g., Melson and Thompson, 1971; Blanchard et al., 1976). Such rocks may be formed by the floating of plagioclase at some depth and the crystallization of euhedral olivine in a shallower environment.

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