

### 53. PETROLOGIC EVOLUTION OF THE MARIANA ARC AND BACK-ARC BASIN SYSTEM— A SYNTHESIS OF DRILLING RESULTS IN THE SOUTH PHILIPPINE SEA<sup>1</sup>

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#### ABSTRACT

The results of Deep Sea Drilling Project Leg 60 pertaining to the petrologic evolution of the Mariana Trough and Mariana arc are summarized in Part I. The rocks recovered at five principal sites are exceptionally diverse, including gabbros and calc-alkalic andesites derived from the West Mariana Ridge; fresh and hydrothermally altered basalts in the Mariana Trough; and interbedded boninites and arc tholeiites in the forearc region between the modern arc and the trench. Part I outlines the stratigraphy, alteration, and petrology of these rock suites and summarizes hypotheses on their origin.

Part II integrates these results with those of previous DSDP legs in the region, and with island sampling, to present a composite 40-m.y. history of the Mariana arc system. Arc volcanism began in the Eocene, probably as a result of a change to the west in the motion of the Pacific plate. A north-south-trending fracture zone was transformed into a trench-subduction complex, trapping a portion of the Pacific plate in the Philippine Sea. The crust of this basin, drilled in Hole 447 (Leg 59), has the composition of typical ocean floor basalt.

The earliest Eocene arc was built up dominantly of arc tholeiite and boninitic lavas, with lesser calc-alkalic lavas, based on the results of Leg 60 drilling at Sites 458 and 459 in the forearc region; Leg 59, Site 448, on the Palau-Kyushu Ridge; and exposures on the islands of Palau, Guam, and Saipan. Near Sites 458 and 459, the forearc crust is thin, formed entirely under water, and includes no known component of ocean crust. Nevertheless it has many of the features of an ophiolite, produced *in situ* by earliest arc volcanism. In the Miocene, the Palau-Kyushu Ridge was split from this ancestral arc by opening of the Parece Vela back-arc basin. A new arc formed on the east side of this basin, the West Mariana Ridge. Lavas recovered from this ridge at Sites 451 and 453, and dating from this time on Guam, have calc-alkalic compositions. In the Pliocene, a second back-arc basin formed, the Mariana Trough, splitting off the West Mariana Ridge and causing volcanism along it to cease. The present arc has been built again on the east side of the new back-arc basin. Based on the compositions of ash at Site 453 and island exposures, the new arc first erupted arc tholeiites but has shifted to calc-alkalic compositions in the past 200,000 years. Both the Parece Vela Basin and Mariana Trough contain tholeiitic basalts with most of the features of mid-ocean-ridge tholeiites, but basalts in the Trough have some of the trace element characteristics of island arc basalts.

Petrogenetic models for the evolution of the arc system therefore must provide for four distinctive magma types—boninites, arc tholeiites, calc-alkalic lavas, and back-arc-basin tholeiites—and explain their distribution in time and space. The relative contributions to the production of these magma types of the subducted slab, the overlying mantle wedge composition, partial melting, and fractional crystallization need to be considered. One important conclusion is that segregation of cumulus gabbros at elevated  $p(\text{H}_2\text{O})$  such as those cored at Site 453 may be a major mechanism for producing calc-alkalic andesites in the Mariana arc system. Immediately following back-arc rifting, when sources of magma are quite shallow and heat flow is high, the conditions necessary for such fractionation would be unlikely to occur. Thus following each episode of back-arc basin rifting, arc volcanism should be tholeiitic. Only later will it become calc-alkalic. We also propose that progressive changes in the mantle beneath each arc resulting from voluminous extraction of basalts may contribute to conditions necessary for new episodes of back-arc rifting.

#### GENERAL INTRODUCTION

Deep Sea Drilling Project Leg 60 completed the transect of the Mariana arc-interarc basin system in the Philippine Sea, begun on Leg 59. Between them the two cruises successfully recovered basement samples from all the main physiographic features in the area (Fig. 1): the West Philippine Sea (Site 457), the Palau-Kyushu Ridge (Site 448), the Parece Vela Basin (Sites 449 and 450), the West Mariana Ridge (Sites 451 and 453), the Mariana Trough (Sites 454 and 456), the Mariana Ridge (Site 457), the Mariana forearc (Sites 458 and 459), and the wall of the Mariana Trench (Sites 460 and 461). Sites 453 through 461 were drilled during Leg 60. Additionally, Leg 58 drilled basement samples from the Shikoku

Basin (Sites 442, 443, and 444), which is a northward continuation of the Parece Vela Basin and roughly equivalent to it in age, and from the Daito Basin in the northern West Philippine Sea (Site 446). The petrological and geochemical information now available for these samples, together with that assembled as a result of field investigations and dredging operations, provides a body of data more comprehensive than that available for any other arc-marginal basin system.

We have divided this chapter into two parts. In Part I, we summarize the results of basement drilling during Leg 60, drawing together principal results from all contributions to this volume dealing with basement samples. These results concern not only primary petrogenesis, but basement structure and alteration as well. In Part II, we review the primary geochemical data from Legs 58, 59, and 60 in relation to possible models for the petrological evolution of island arc-marginal basin sys-

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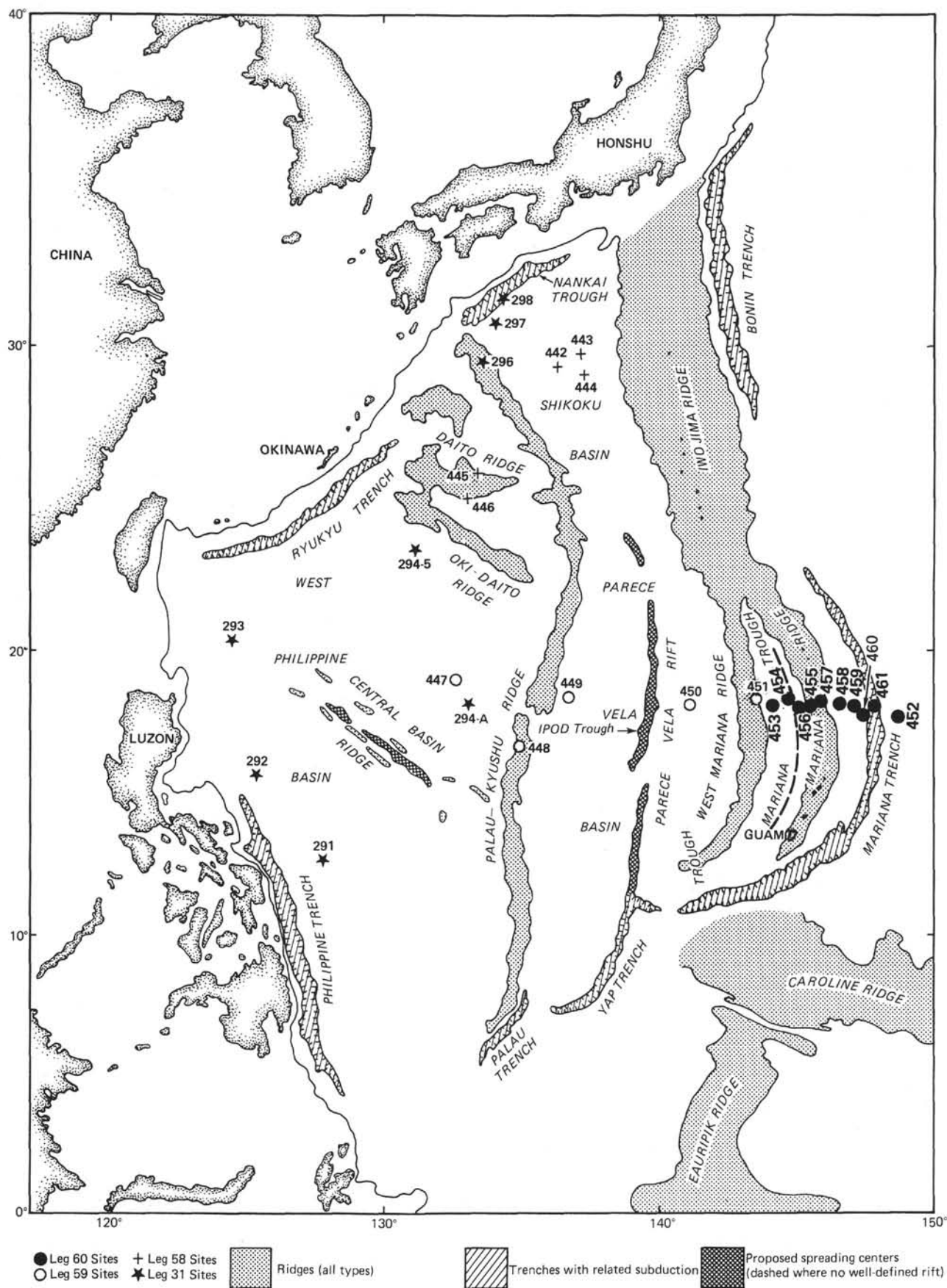


Figure 1. Location of DSDP sites drilled in the Philippine Sea. Site survey targets (SP numbers) are also indicated.

tems and for the Mariana-Bonin arc system in particular. A considerable body of geophysical and tectonic data is already available for the Philippine Sea and forms a valuable background on which to construct petrological-geochemical models for the evolution of intra-oceanic island arcs. In addition, the sedimentary history of the drill sites provides information about the timing and patterns of volcanism and interarc basin rifting, refining considerably the original interpretations of Karig (1971a, 1971b, 1975). These will be the principal concerns of this report.

## PART I. SUMMARY OF PETROLOGICAL INVESTIGATIONS BASED ON LEG 60 DRILLING

### Leg 60 Basement Drilling Data

Site locations, water depth, depth of basement, and recovery data for basement drilled during Leg 60 are given in Table 1. Altogether, basement drilling was very successful. Over 600 meters total basement penetration was obtained at five principal sites, and a variety of rock types were recovered. The following sections summarize the principal results pertaining to igneous basement for each site. Where applicable, the various lithologic, chemical, and magnetic units are defined, and a summary of the principal petrological and geochemical features of the rocks, as well as the state of their alteration, is given. Relevant geophysical information from site and other surveys and from shipboard programs (mainly heat flow) is considered. Additional interpretation of regional tectonic history in light of the drilling is in the synthesis by Hussong and Uyeda (this volume).

### Drilling in the Mariana Trough

Four sites (453–456) were drilled in the Mariana Trough (Fig. 1) in an effort to find a location where a deep penetration multiple reentry hole could be drilled and to establish the spreading history of this youngest Mariana back-arc basin. Although no reentry site was found, igneous and metamorphic rocks were recovered in four holes at three sites. Site 453 was located in a small north-trending sediment pond on the western side of the Trough. A complex gabbro-metabasalt polymict breccia was cored for 146 meters beneath 450 meters of turbiditic vitric tuffs ranging in age from late Pliocene to Recent. Pillow basalts underlying Pleistocene sediments were cored nearer the center of the Trough in

Holes 454A, 456, and 456A. The latter two holes were about 200 meters apart. In each of these holes, basement penetration was low, and hole instabilities precluded reentry. Detailed operational, lithologic, and petrographic summaries of the rocks recovered in all these holes can be found in the site chapters.

### Polymict Breccias at Site 453

Three types of igneous and metamorphic rock sequences were recovered in Hole 453. The upper two are breccias containing coarse cobbles or boulders of igneous and metaigneous rocks (Fig. 2). The deepest is a sheared and highly altered gabbro, recovered in two fragments totalling 150 cm in length. Gabbros and metabasalts occur in the upper breccia (86 m thick) with rare minor lithologies, including both foliated and non-foliated quartz diorite (Sharaskin, this volume). Some gabbro clasts reach 40 cm in longest cored diameter, and where such large boulders are abundant, core recovery is highest. Metabasalts are either aphyric or have large recrystallized plagioclase phenocrysts. There are no large metabasalt boulders, although cores contain intervals where several adjacent metabasalt pieces evidently came from the same boulder.

The lower breccia contains aphyric greenstones, some quite pyritized, with the compositions of metamorphosed andesites and dacites (Wood et al., this volume). The rocks are small (3–7 cm), pale green, and are associated with a small amount of indurated pyroclastic breccia, also metamorphosed, and soft, dark green, friable breccia, perhaps originally a hyaloclastite. Some of the lithic fragments are veined with epidote.

The rocks have experienced a variety of low- and high-grade metamorphic events. The highest grade of metamorphism is represented by gabbros containing anorthite, green aluminous amphibole, and green hercynitic spinel (see Natland, gabbros chapter, this volume). Some gabbros have been completely recrystallized to prehnite-clinzoisite-plagioclase assemblages. Others have been partially recrystallized to chlorite and actinolite, with almost complete transformation of mafic minerals but only partial sericitization of plagioclase. Still other gabbros shows tectonized fabrics, including extensive development of (100) twinning on clinopyroxene, and bent plagioclase crystals. Greenschist facies metamorphic conditions evidently also affected the metabasalts, which contain groundmass andesine-oligoclase,

Table 1. Basement drilling and recovery data, Leg 60.

Hole	Location		Water Depth (m)	Depth to Basement from Seafloor	Basement Cored (m)	Basement Recovered (m)	Basement Recovered (%)
453	17°40.17'N	143°40.95'E	4693	455.5	149.50	33.30	22.3
454A	18°00.78'N	144°31.92'E	3819	66.0	104.50 <sup>a</sup>	19.95	19.1
456	17°54.68'N	145°10.77'E	3590.5	134.0	35.00	3.66	10.5
456A	17°54.71'N	145°10.88'E	3591	118.5	40.50	2.34	5.8
458	17°51.84'N	146°56.06'E	3449	256.5	209.00	35.79	17.1
459B	17°51.75'N	147°18.09'E	4115	558.5	133.00	24.12	18.1
Totals					671.50	119.16	17.7

Note: Basement in Hole 453 is gabbro-metabasalt polymict breccia. At all other sites it is pillow lavas. Igneous and metamorphic rocks were also recovered as talus in Holes 460 (17°40.14'N; 147°37.92'E), 460A (17°40.02'N; 147°35.16'E), 461 (17°46.05'N; 147°41.18'E), and 461A (17°46.01'N; 147°41.26'E), in water depths between 6443 and 7034 m in the Mariana Trench.

<sup>a</sup> Includes sediments interbedded with basalts.

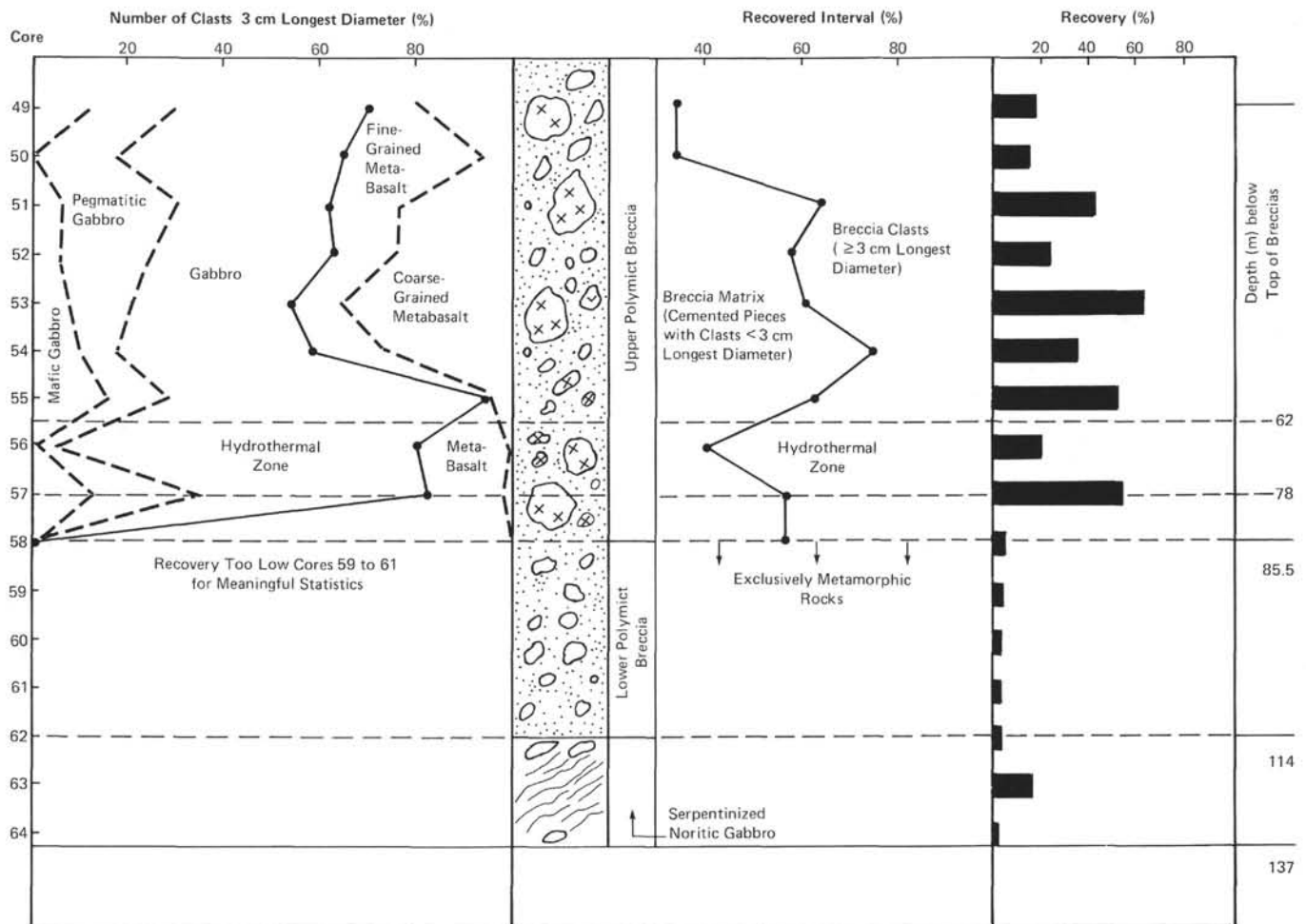


Figure 2. Stratigraphy, schematic lithology, and recovery per core of gabbro polymict breccias cored in Hole 453.

amphibole, chlorite, epidote, clinozoisite, and sphene. They are not foliated but have granoblastic and porphyroblastic textures. Some basalts of similar compositions experienced retrograde(?) prehnite-pumpellyite facies metamorphism (Sharaskin, this volume).

Despite the evidently high grades of metamorphism and the probability that the metabasalts were metamorphosed prior to emplacement in the breccias, much of the metamorphism was hydrothermal and occurred after the breccias were deposited. The alteration is concentrated in two zones lower in the breccias (Fig. 2). Its effect was to form chlorite-actinolite aggregates in some rocks and to riddle many of them with chlorite veins. The breccia matrix, which originally consisted of crushed angular fragments and chips from the larger clasts, has been intensely altered, especially in the two zones just mentioned. Mafic minerals throughout most of the upper breccia matrix (originally olivine, diopside, amphibole, and biotite) are now largely or entirely altered to serpentine, talc(?) chlorite, and abundant reddish iron hydroxides. In the zones of intense alteration, the matrix is green, not red, in color and is almost completely chloritized. One effect of chlorite formation has been to redistribute potassium, which does not enter the chlorite crystal lattice. It was originally fairly abundant in such minerals as amphibole and biotite in the breccia matrix

and in the more evolved rock types (andesite, dacite) in the lower breccias. Additional potassium probably was provided by seawater. The potassium now occurs in almost pure potassium feldspar in the chloritized breccias, replacing calcic plagioclases around the edges of coarse gabbro cobbles and completely replacing many plagioclase chips in the matrix, giving them a bleached appearance. In the zone of intense alteration in the upper breccia (Cores 55 and 56, Fig. 2), this alteration has completely demagnetized the gabbros (Bleil, this volume). Both the formation of chlorite and the exceeding of Curie temperatures of the gabbros imply quite elevated temperatures of hydrothermal alteration. Oxygen isotope data obtained on drusy quartz in the topmost piece of breccia recovered, and on potassium feldspars, indicate temperatures of formation of these minerals of about 100°C near the top of the breccias and 200°C in the highly chloritized zone (Lawrence and Natland, this volume). The association of chlorite with epidote suggests that temperatures locally exceeded 250°C, based on analogy with marine geothermal systems (e.g., Humphris and Thompson, 1978; Mottl and Holland, 1978) and Iceland (Tomasson and Kristmannsdottir, 1972). The occurrence of minerals such as prehnite, actinolite, chlorite, and biotite suggests that temperatures may even have reached 350°C, based on comparisons with



the Cerro Prieto and Salton Sea geothermal systems (Elders et al., 1979; McDowell and Elders, 1980). These greenschist facies assemblages did not necessarily form at elevated pressures, as the Salton Sea and Cerro Prieto occurrences demonstrate. Those gabbros with aluminous amphibole and green spinel apparently indicate elevated pressures (Natland, gabbros chapter, this volume), but these minerals clearly formed before the breccias were emplaced and hydrothermal alteration occurred.

Some alteration, but not at quite these temperatures, may be occurring today. A heat flow of 2.1 heat flow units (HFU), or 88 mW/m<sup>2</sup>, was measured in Hole 453 (Uyeda and Horai, this volume). This extrapolates to temperatures at the top of the breccias of about 50°. Moreover, sediment pore fluids show sharp reversals of Mg<sup>2+</sup> and Ca<sup>2+</sup> just above the breccias to near seawater compositions. Gieskes and Johnson (this volume) infer from this that little-modified seawater must be flowing beneath the sediments at the top of the still-permeable breccias, driven by hydrothermal convection. These must have flushed out pore fluids produced during extensive chloritization of the breccias, which should have had compositions considerably different from seawater (i.e., lower Mg<sup>2+</sup> and higher Ca<sup>2+</sup>).

These various types of metamorphism and alteration modified the compositions of the volcanic rocks and gabbros in the breccias. The most important effects are removal of CaO and probable addition of MgO during higher grades of metamorphism and addition of K, Rb, and Ba during chloritization of the breccia matrix and transformation of plagioclases to K-feldspar. These effects result in a higher proportion of normative albite, lower normative hypersthene (or more olivine), and higher normative orthoclase, respectively, than in little altered basalts of similar composition recovered at DSDP Site 451, on the West Mariana Ridge, 44 km away (Natland, gabbros chapter, this volume). The hydrothermally altered metaandesites and dacites of the lower breccia have been almost entirely albitized compared with fresh andesites and dacites from the Mariana arc system.

Despite these alteration effects, the original gabbros and volcanic rocks now in the breccias are undeniably calc-alkalic (Wood et al., this volume). Alteration has not removed the signature of fairly low iron enrichment that characterizes calc-alkalic suites, and, more important, much of the calc-alkalic trace element chemistry of the rocks is intact. Elevated abundances of Sr and light rare-earth elements show them to be kindred to little-altered calc-alkalic rocks from the Mariana arc (Dixon and Batiza, 1979) and the West Mariana Ridge (Mattey et al., 1980; Wood et al., 1980).

The gabbros have a remarkable and surprisingly restricted range of compositions expressed in terms of "cryptic variation" parameters such as Mg/Mg + Fe and normative An/Ab + An. Of 31 gabbro samples analysed from this site, all but 4 have normative An/(An + Ab) > 0.75, and in many this ratio exceeds 0.90 (Fig. 3A). Most conform petrographically to olivine gabbro, but with widely varying modal proportions of plagioclase and clinopyroxene. A few gabbro samples

are olivine-free and consist of clinopyroxene, plagioclase, biotite, magnetite, and ilmenite or of amphibole and plagioclase, with minor opaques. Most of the samples have heteradcumulus textures, but others have been recrystallized and are now granoblastic. The plagioclase in olivine gabbros is extremely calcic, and anorthite (An<sub>94-97</sub>) in composition. Olivines are fairly iron-rich (Fo<sub>72-76</sub>), clinopyroxenes are diopsidic to salitic, and amphiboles are hornblendic except in the metamorphosed gabbros (Natland, gabbros chapter, this volume). The metagabbros with green hercynitic spinel are few and are compositionally similar to the olivine gabbros. Their spinel and aluminous amphiboles appear to represent a subsolidus reaction between olivine and plagioclase in the presence of water.

The calcic plagioclase and high normative anorthite contents of the whole-rock analyses distinguish these gabbros from gabbros in fracture zones, in ophiolites, and in stratiform complexes such as the Bushveld or Skaergaard (Fig. 3A). They most resemble gabbros occurring as ejected blocks from Soufrière volcano, St. Vincent, West Indies (Lewis, 1973; Arculus and Wills, 1980); as zoned fragments of plutons tectonically emplaced at high levels among granodioritic and quartz-dioritic plutons of the Peninsular Range batholith in Southern California (Nishimori, 1976); and as xenoliths in lavas from Agrigan volcano in the Mariana Arc (Stern, 1979), all calc-alkalic associations. The average composition of the gabbros at Site 453 is nearly identical to the composition of a cumulus gabbro mass calculated by Stern to underlie Agrigan and to be responsible for its calc-alkalic trend (Fig. 3B).

To explain the compositions and mineralogy of the similar Peninsular Range gabbros, Nishimori (1976) argued that  $p(\text{H}_2\text{O})$  of up to 2 kbar caused the liquidus to shift to more calcic plagioclase compositions (following arguments of Yoder and Tilley, 1962; and Yoder, 1969). Fractionation of mineral assemblages with a high proportion of such calcic plagioclase (which has only about 43% SiO<sub>2</sub>) would cause rapid SiO<sub>2</sub> enrichment in residual liquids before there could be much iron enrichment. Such fractionation involving removal of material equivalent to the Site 453 gabbros could explain the calc-alkalic variation of West Mariana Ridge lavas cored at Site 451, 44 km to the west (Mattey et al., 1980).

The gabbros are proof of the sundering of an island arc during back-arc spreading. They also indicate that a portion of this Miocene arc must occur on the eastern side of the Mariana Trough—which is evident from coeval volcanic exposures on Guam (Stark, 1963). If the high grades of greenschist facies metamorphism—particularly the formation of aluminous amphibole and green spinel—manifest in the gabbros and metabasalts indicate elevated pressures, then considerable uplift (perhaps 10 km or more) attended rifting. This seems excessive in the context of simple rifting. A hydrothermal origin for the metamorphic minerals eases this problem considerably. Still, some uplift must have occurred to have caused the exposure of gabbros which crystallized at elevated  $p(\text{H}_2\text{O})$ . There is at present about 4.5 km of relief on the West Mariana Ridge, more than half the to-

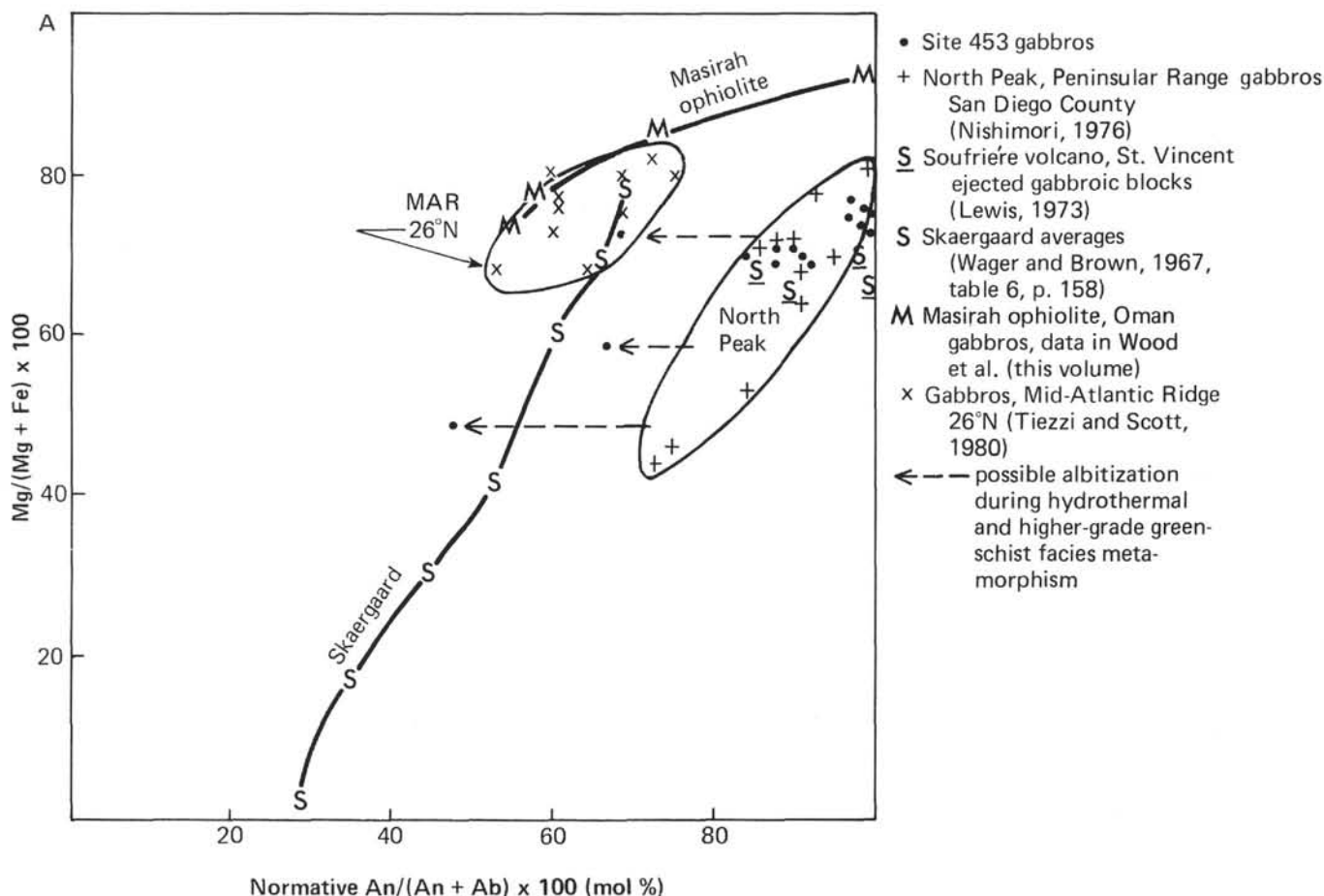


Figure 3. A.  $\text{Mg}/(\text{Mg} + \text{Fe}) \times 100$  versus normative  $\text{An}/(\text{An} + \text{Ab}) \times 100$  for calc-alkalic gabbros compared with ocean crust, ophiolitic, and layered gabbro intrusions. B.  $\text{Mg}/(\text{Mg} + \text{Fe}) \times 100$  versus normative  $\text{An}/(\text{An} + \text{Ab}) \times 100$  for West Mariana Ridge and Agrigan Island, Mariana arc, compared with fields for Site 453 and Peninsular Range, southern California batholith (Nishimori, 1976), and the average cumulus mass beneath Agrigan calculated by Stern (1979) to be responsible for the Agrigan fractionation trend (solid line). The effect of albitization on Hole 453 lower breccia metavolcanic samples is indicated by dashed arrows.

total amount of combined exhumation and uplift needed to expose gabbros which crystallized at  $p(\text{H}_2\text{O}) \approx p$  (total) of about 2 kbar.

The nearest sources for the breccias are the walls of the sediment pond in which Site 453 was cored. The closest and most likely source is a small promontory or spur to the northwest which appears to link directly with the West Mariana Ridge (Packham and Williams, this volume). Based primarily on the flatness of the breccia surface in the sediment pond as seen in airgun profiles (Site 453 chapter, this volume), Natland (gabbros chapter, this volume) and Sharaskin (this volume) believe that the breccias most closely resemble deposits of catastrophic rock falls or "stürztstroms" (Hsü, 1975). Alternatively, "fans" of mass-wasted volcanic and plutonic rocks with slopes as little as 2° have been seen in large North Atlantic transform-fault zones (P. J. Fox, personal communication). The process of breccia emplacement at Site 453, then, need not have been catastrophic. The breccia formation process ceased completely sometime prior to 5.0–5.2 Ma, the time of formation of the oldest paleontologically dated turbidite at Site 453 (Ellis, this volume).

In summary, the occurrence of polymict breccias at Site 453 is the result of two fundamentally different thermal-tectonic events. The rocks represent both the extrusive and intrusive products of calc-alkalic crystal fractionation at elevated water pressure. The exposure of these rocks and their chaotic dismemberment into breccias was caused by back-arc basin rifting prior to 4.5 Ma. The rifting apparently reached the very heart of the arc, reaching solidified magma chambers deep within it. The hydrothermal alteration of portions of the breccias is clearly related to geothermal activity in the Mariana Trough. Whether this was caused by magmatic injection beneath Site 453 cannot be confirmed by present data, although the high temperatures of hydrothermal alteration probably indicate a magmatic source of heat nearby.

#### Interarc Basin Basalts at Sites 454 and 456

The lithology, distribution of chemical types, and location of changes in magnetic inclinations downhole are shown in Figure 4 for the three holes penetrating basalts at Sites 454 and 456, located on either side of, yet close to the center of spreading in, the Mariana Trough. The

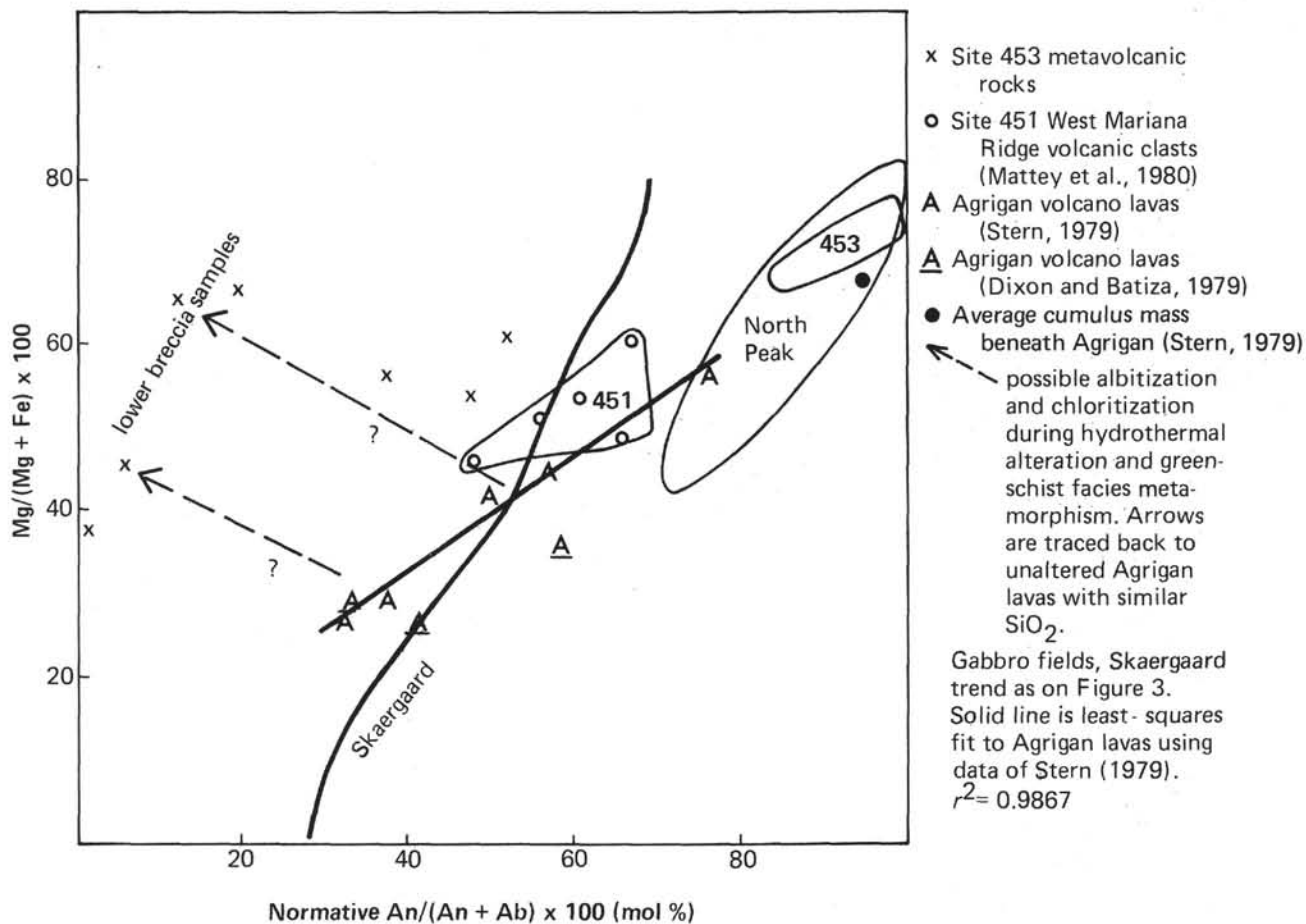


Figure 3. (Continued).

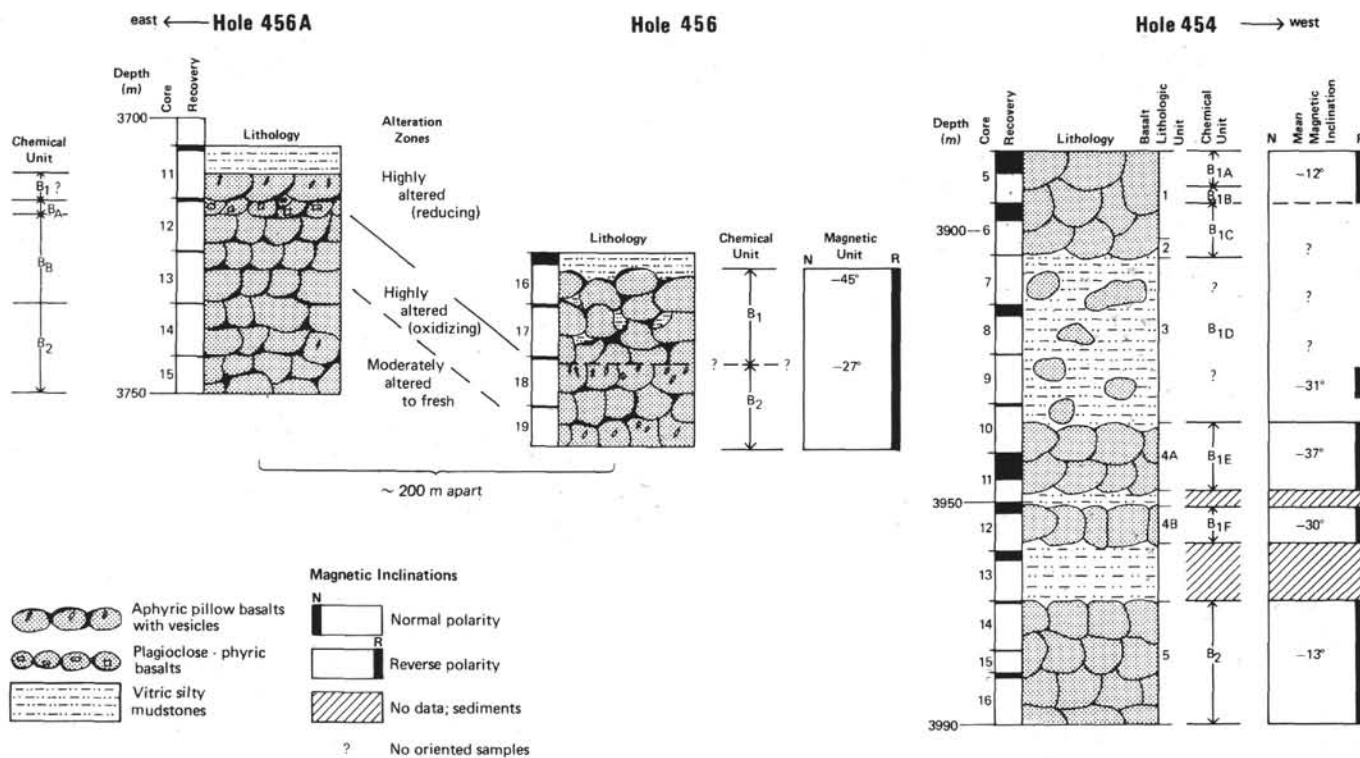


Figure 4. Basement lithologic, chemical, and magnetic stratigraphy, Sites 454 and 456, Mariana Trough.

oldest sediments at Site 454 are about 1.6 m.y. old and at Site 456 are 1.8 m.y. old (Kling, this volume; Ellis, this volume). The basalts at Site 454 are interbedded with vitric mud. The two holes at Site 456 are about 200 meters apart and there are no interbedded sediments. However, one basalt piece in Hole 456 has a recrystallized sedimentary inclusion several centimeters in diameter. The basalts of both sites are mostly sparsely phyrlic or microphyric vesicular pillows and flows, moderately to highly olivine-rich. Only one basalt type in Hole 456A was strongly plagioclase and olivine phyrlic (Fig. 4). The basalts are quite vesicular, although great depth of eruption (~3500 m) has undoubtedly kept the size of vesicles quite small. The upper massive flow in the middle of the basement section at Site 454 is more vesicular, hence less dense, toward the top, a feature readily revealed by the density logs (Fig. 5). At Site 456, many vesicles are filled with secondary minerals, including quartz, calcite, and pyrite, giving the rocks higher than usual thermal conductivities (Horai, this volume). The

greater abundance of vesicles in these basalts than in normal mid-ocean-ridge basalts (MORB) is suggestive of a higher primary volatile content (cf. Garcia et al., 1979; Muenow et al., 1980), though this is not necessarily typical of all interarc basin basalts.

Several chemical types occur at Sites 454 and 456, identified by the letter B (for basalt) and an alphabetical or numerical subscript on Figure 4. The various types of basalt can be distinguished by trace element variation, such as that of Ni versus Zr (Fig. 6). Those chemical types present in Hole 456A but not Hole 456 are given alphabetical subscripts. The basalts have the major element characteristics of little-fractionated or moderately fractionated mid-ocean-ridge basalts (i.e.,  $\text{TiO}_2 \approx 1.0\%$ ;  $\text{K}_2\text{O} \approx 0.3\%$ ;  $\text{Mg}/\text{Mg} + \text{Fe} = 0.6\text{--}0.7$ ). Despite this, only a few of them have the geochemical characteristics of depleted normal (N-type) MORB. Wood et al. (this volume) note in particular "higher Sr, Ba, Th, and light REE contents relative to Zr, Ti, Y, and the heavy REE than N-type MORB." They also are depleted in Ta and

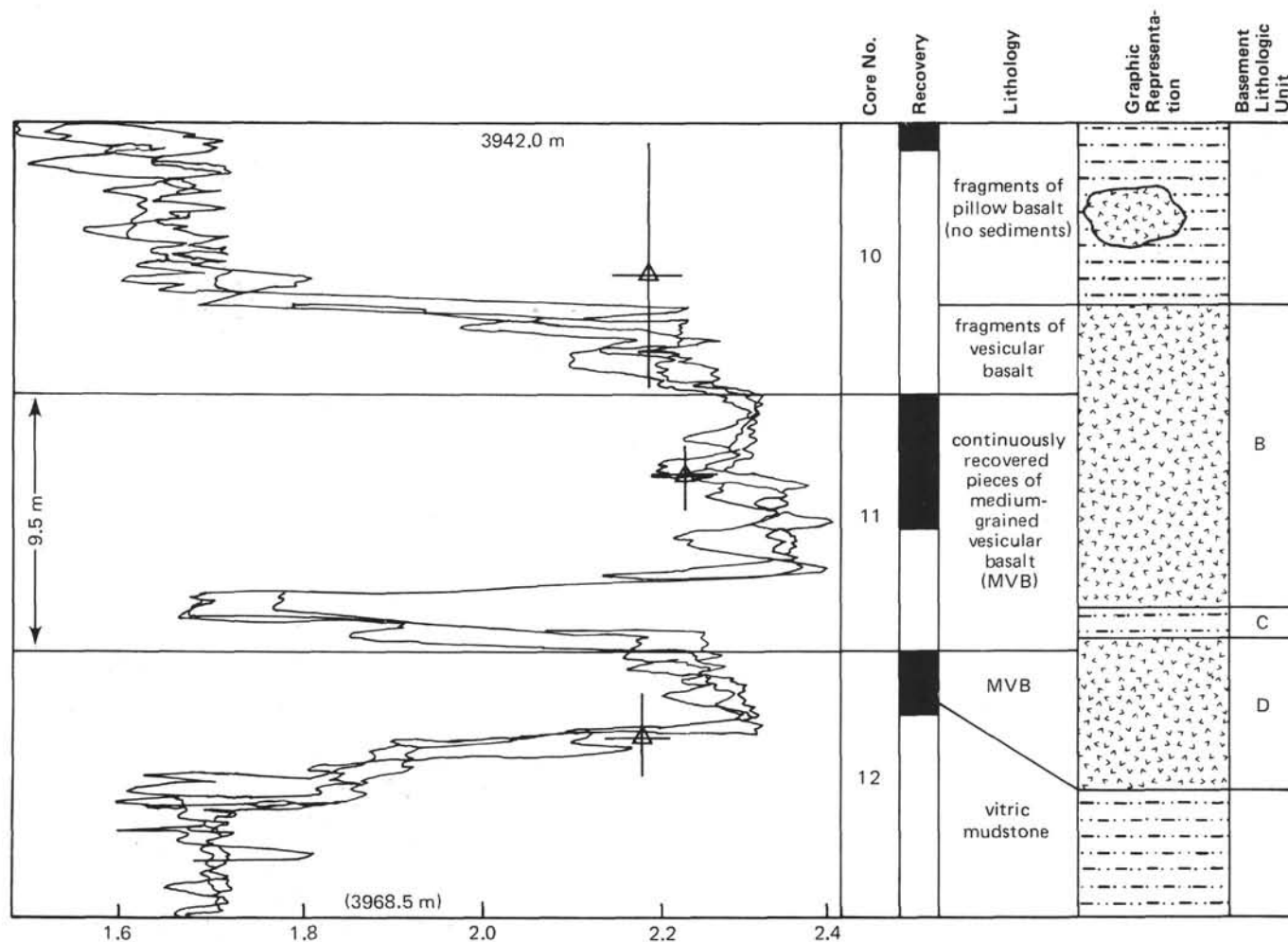


Figure 5. A portion of the three density logs (overlapped) obtained in Hole 454A (Mariana Trough) during Leg 60, showing the inferred relationship to interbedded massive basalts, pillow basalts, and sediments. Core number, recovery, and lithologic units are also shown. Triangles with error bars indicate locations of samples and shipboard density measurements obtained on them by T. Francis. Note decrease in density upward in upper massive basalt, caused by vesicles.



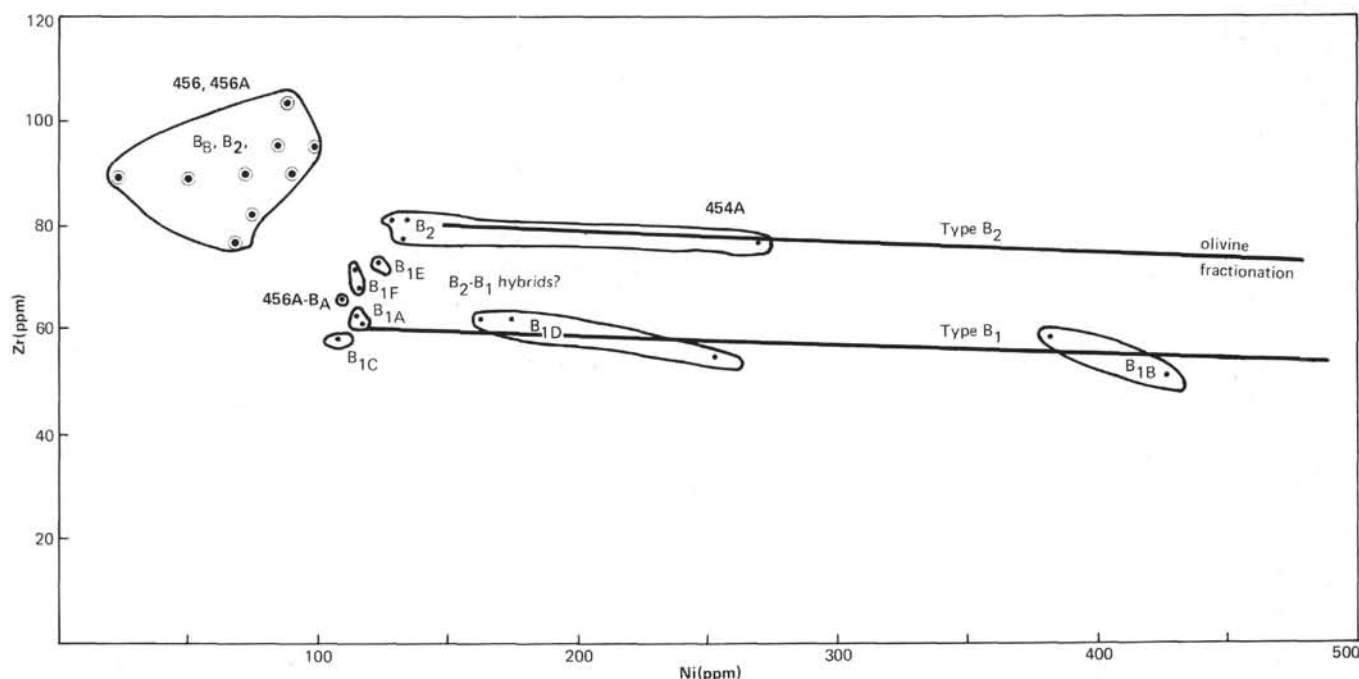


Figure 6. Ni versus Zr for basalts of Holes 454A, 456, and 456A, Mariana Trough. (• = Hole 454A; ○ = Holes 456 and 456A. Chemical types and subtypes of Figure 4 are indicated. Data from Wood et al., this volume.) In Hole 454A, olivine-controlled Type B<sub>2</sub> was succeeded by olivine-controlled B<sub>1</sub> subtypes with systematically lower Zr. Subtypes B<sub>1E</sub> and B<sub>1F</sub>, which stratigraphically represent the earliest B<sub>1</sub> subtypes, have intermediate Zr, suggesting that they may be hybrids. Ni concentrations in any given sample are probably controlled by local accumulations of olivine in pillows and flows. Holes 456 and 456A samples are more fractionated than those from Hole 454A; Ni abundance probably was affected by alteration.

Nb relative to Th and La. This geochemical signature is characteristic of island arc basalts (Saunders et al., 1980). Such features seem to be gradational throughout the basalts of Hole 454A and occur in one unit (the plagioclase-olivine strongly phyrlic basalt type B<sub>A</sub>) of Hole 456A. Wood et al. (this volume) argue that this could represent mixing of depleted N-type MORB with island arc magmas at Site 454 and interbedding of the two magma types at Site 456. It could also reflect modification of the mantle source by fluids carrying large-ion lithophile (LIL) elements which were derived from dehydration of the subducted slab or from sediments carried down the subduction zone. Alternatively, the anomalous chemistry might be produced by assimilation of the intimately associated glass-rich sediments, derived from the Mariana arc, and which would have the required "arc" component. However, the fact that similar geochemical features are apparent in other marginal basin basalts, at least during the early stages of back-arc spreading (Tarney et al., 1981), lends support to the suggestion that the cause is more fundamental and related to modification of the mantle source composition prior to back-arc spreading.

There is good trace element evidence that the various chemical types are not related by crystal fractionation before or after possible contamination. As an example, substantially higher Zr levels occur in basalt type B<sub>2</sub> at Site 454 than in basalt type B<sub>A</sub>, despite similar ranges of Ni (Fig. 6). Ni abundance in these samples is controlled primarily by olivine fractionation, which produces the parallel trends on Figure 6. The trends are an indication

that a range of parental compositions are produced either by variable degrees of melting or by differences in the abundances of low-partition-coefficient (hygromagmatophile, or HYG) elements in the mantle source of these basalts. Such variations resemble those along certain portions of the Mid-Atlantic Ridge (e.g., DSDP Sites 332 and 395; Natland, 1978; Sites 409–413, Tarney et al., 1979).

Fryer et al. (this volume) argue that basaltic glasses from Sites 454 and 456, as well as those from dredge hauls from the central axis of the Mariana Trough, resemble other interarc basin basalts, particularly those of the Scotia Sea (e.g., Saunders and Tarney, 1979), but differ significantly in major element composition from normal MORB. The features they single out are (1) generally lower TiO<sub>2</sub>, (2) generally higher Na<sub>2</sub>O, (3) lower FeO\* (total iron as FeO), and (4) higher Al<sub>2</sub>O<sub>3</sub> when specified at a given MgO content. They also note some of the trace element differences cited by Wood et al. (this volume), somewhat higher <sup>87</sup>Sr/<sup>86</sup>Sr (Hart et al., 1972) as well as more H<sub>2</sub>O-rich volatile inclusions (Garcia et al., 1979) than in MORB.

Alteration at Site 454 was at low temperatures and oxidative, producing brown or reddish alteration rinds on some basalts but generally having little effect on major oxide chemistry. At Site 456, however, alteration was both extensive and diverse, resulting in marked changes in the bulk compositions of the basalts. In both Holes 456 and 456A, at the top of basement, alteration was nonoxidative and occurred at high temperature, resulting in the large-scale development of chlorite in the

basalts and of pyrite, opal, quartz, and chalcopyrite (Natland and Hekinian, this volume). Recrystallized tuffs just above basalt contain quartz and wairakite, which typically forms under high temperature hydrothermal conditions in silica-rich materials and has been produced experimentally only above 250°C (Liou, 1971).

Below the chloritic, pyritized basalts, over a vertical distance of only a few meters, alteration becomes distinctly oxidative in both holes, with extensive development of reddish alteration rinds and of soft, friable, clay-rich basalts. At the bottom of each hole, even this alteration disappears and fresh glass occurs.

The chemical effects of the alteration contrast in the nonoxidative and oxidative zones. In the nonoxidative zone, there is considerable variation in CaO and MgO among basalts within distinct chemical types (defined on the basis of immobile oxides such as TiO<sub>2</sub> and trace elements such as Zr). Yet there is virtually no variation in very low abundances of K<sub>2</sub>O. In the zones of oxidative alteration, however, K<sub>2</sub>O, Rb, and Ba vary considerably from sample to sample, evidently reflecting the formation of abundant K-Fe clay minerals in the samples.

The simple downhole sequence of nonoxidative and oxidative zones of alteration followed by fresh basalts suggests that one episode of alteration was involved in which high-temperature nonoxidative hydrothermal fluids made their way along the permeable sediment/basalt contact (Natland and Hekinian, this volume). The zone of oxidative alteration was a zone of intense mixing with oxygenated pore fluids (essentially seawater). K, Rb, or Ba leached from basement rocks and carried by the high-temperature fluids did not enter the chlorite crystal lattices forming in the nonoxidative zone, but continued flux and mixing of such fluids with seawater in the oxidative zone favored formation of K- and Fe-rich clay minerals. Below the mixing zone, neither the temperature of the fluids nor their diluted composition produced significant alteration.

### Fore-arc and Trench Drilling

The objective of drilling in the fore-arc region and the trench was to determine if fragments of ocean crust have been emplaced by faulting into the landward wall of the Mariana Trench. The principal result of this drilling was to establish quite the contrary. In two holes in the fore-arc region, 458 and 459B, extrusive igneous basement was reached beneath arc-produced Eocene-late Oligocene turbidites. In Hole 458, the lavas include an unusual rock type, boninite, known elsewhere only from island arc settings and certain ophiolites. These are interbedded with island arc tholeiites, which also were recovered in Hole 459B. Although there were drilling difficulties in igneous and metamorphic talus in four holes at two sites in the trench (460, 460A, 461, and 461A), those rocks, too, have primarily the compositions of island arc tholeiites and in Hole 460 include a boninite. No ocean crust or marine pelagic sediments were cored in any of these holes.

The following sections provide an evaluation of lava stratigraphy in Holes 458 and 459B and a summary of

the principal hypotheses presented in this volume concerning the origin of the fore-arc lavas and structure. Operational summaries and detailed lithologic and petrographic descriptions are in the site chapters.

### Igneous Rock Series in the Fore-arc and Trench Region

The boninitic rocks of Hole 458 are the first such lavas ever recovered by drilling. Although boninites are unusual, they have chemical and petrographic characteristics that are of great importance to the understanding of island arc magmatism. The boninites of Hole 458 are of particular importance because island arc tholeiites occur in the same hole and because we can specify that they erupted in late Eocene-early Oligocene times, early in the history of the Mariana arc. Outside this volume, the Hole 458 boninites have already been repeatedly cited with reference to the earliest stages of arc magmatism (Cameron et al., 1979, 1980; Beccaluva et al., 1980; Meijer, 1980) and because of their relevance to ophiolites (Cameron et al., 1979, 1980). Here, we focus on the stratigraphy of the boninites and arc tholeiites of Hole 458 and 459B and consider possible geochemical relationships among the rock types. Other chapters in this volume may be consulted for details of petrography and mineralogy (Meijer et al. and Natland, crystal morphologies chapter), geochemistry (Wood et al., Bougault et al., and Hickey and Frey), and experimental petrology (Kushiro).

The chemical and magnetic stratigraphy of volcanic basement in Holes 458 and 459B is shown in Figure 7. In both holes, the rocks consist of alternating pillows and more massive flows or possibly intrusives. In each hole, several distinctive chemical and magnetic units were identified. Chemical units are defined on the basis of one or more essentially identical, or at least very similar, chemical analyses in stratigraphic order, based mainly on the data of Wood et al. (this volume). In making comparisons, alteration, which affected mainly such mobile elements as K, Rb, and Ba, was taken into account by keying the stratigraphy to less mobile elements such as Ti, Zr, Y, and the like. Also discounted were the effects of possible *in situ* magmatic differentiation, which has occurred in some of the more massive cooling units. Using the data of Wood et al. (this volume), average analyses of the four principal chemical types in Hole 458 (plus comparative individual and average analyses) and seven chemical subtypes in Hole 459B are listed together with trace element averages and CIPW norms in Table 2. There are two principal rock types, designated A and B, referring respectively to boninitic lavas and basalts. The distinct chemical types and subtypes are designated by numerical and alphabetical subscripts on Figure 7.

The boninitic rocks of Hole 458 have been variously referred to in this volume as members of the boninite series (Meijer et al.; Sharaskin), as high-MgO bronzite andesites (Site 458 chapter), or more simply as boninites and bronzite andesites (Natland, crystal morphologies chapter). This lack of consistent nomenclature is a consequence of petrographic dissimilarities between the Hole 458 lavas and certain previously described boninites as

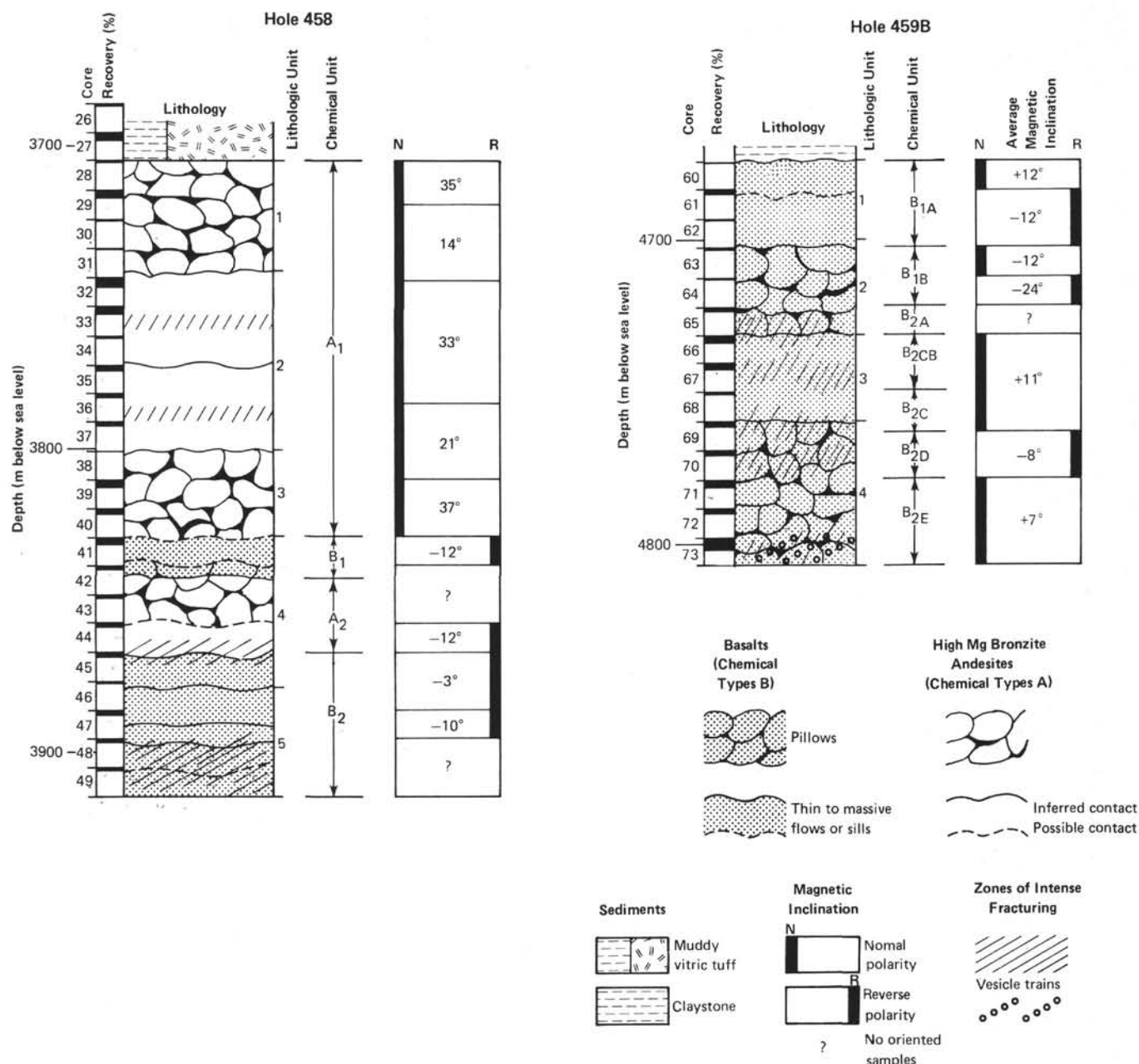


Figure 7. Basement lithologic, chemical, and magnetic stratigraphy, Holes 458 and 459B, Mariana fore-arc region, drilled during Leg 60.

well as of generally lower MgO, Ni, and Cr in Hole 458 samples. For these reasons, Meijer (1980) defines normative criteria to distinguish boninite series lavas from more typical andesites and from basalts. He is critical of early petrographic schemes. Natland (crystal morphologies chapter, this volume) noted that glassy boninites, on which the original petrographic definitions are based, represent only thin outer chilled rims of boninite pillows, and questioned the applicability of the designation "high-MgO" to many of the Hole 458 boninitic lavas. He proposed extending the petrographic definition of boninites to include pillow interiors and massive flows in which plagioclase can be an abundant mineral and, for the purposes of classification, distinguished olivine boninites from simple boninites that lack olivine,

such as the Hole 458 lavas. The lower MgO Hole 458 samples can thus be termed bronzite andesite and included in a boninite series (Bloomer et al., 1979). For convenience we shall use these conventions in the following discussion.

The distinctive petrographic feature of all boninites is crystallization of groundmass microlitic clinopyroxene before plagioclase, even in fairly coarse-grained pillow interiors. The Hole 458 lavas also contain complex bronzite-augite intergrowths in glassy and spherulitic samples, plus separate phenocrysts of both minerals. Temperatures of crystallization of these minerals, calculated using the formula of Ishii (1980), are in the range of  $1250 \pm 60^\circ\text{C}$ , higher than two-pyroxene tholeiitic island arc basalts and calc-alkalic andesites from the



Table 2. Averages of chemical types and subtypes of boninites and island arc tholeiites, DSDP Hole 458 and 459B.<sup>a</sup>

	Hole 458							Hole 459B						
	Ave. Type A <sub>1</sub> Boninite 1(21) <sup>b</sup>	Sample 458-29-2, 37 cm Boninite 2	Ave. of Bronzite Andesite Flow Cores 32-35 3(6) <sup>b</sup>	Sample 458-40-2, 90 cm Boninite 4	Type B <sub>1</sub> Basalt 5 <sup>b</sup>	Ave. Type A <sub>2</sub> Boninite 6(6) <sup>b</sup>	Ave. Type B <sub>2</sub> Basalt 7(6) <sup>b</sup>	Ave. Type B <sub>1A</sub> Basalt 8(3) <sup>b</sup>	Ave. Type B <sub>1B</sub> Basalt 9(2) <sup>b</sup>	B <sub>2A</sub> Basalt 10	Ave. Type B <sub>2B</sub> Basalt 11(3) <sup>b</sup>	Ave. Type B <sub>2C</sub> Basalt 12(2) <sup>b</sup>	Ave. Type B <sub>2D</sub> Basalt 13(2) <sup>b</sup>	Ave. Type B <sub>2E</sub> Basalt 14(4) <sup>b</sup>
SiO <sub>2</sub>	55.0	52.0	58.5	52.8	53.2	53.9	52.0	54.3	51.2	52.6	56.7	57.7	55.7	55.1
TiO <sub>2</sub>	0.32	0.36	0.31	0.35	1.13	0.52	1.10	0.69	0.83	1.21	0.97	0.74	0.91	1.13
Al <sub>2</sub> O <sub>3</sub>	15.2	14.4	14.2	14.9	15.5	14.3	14.0	14.6	14.3	13.9	13.1	13.3	14.3	13.9
Fe <sub>2</sub> O <sub>3</sub>	9.31	10.23	9.12	10.24	9.66	10.54	13.63	9.95	10.67	13.63	11.89	10.82	11.01	11.91
MnO	0.12	0.12	0.11	0.09	0.06	0.13	0.11	0.14	0.14	0.16	0.15	0.12	0.11	0.11
MgO	6.87	9.71	5.62	8.97	3.73	6.21	6.69	5.78	9.64	6.07	5.09	5.52	5.40	5.19
CaO	9.59	8.70	9.79	7.98	6.12	8.33	6.55	10.82	8.47	6.06	7.25	7.63	8.25	7.40
Na <sub>2</sub> O	2.60	2.63	2.48	2.47	5.70	2.65	3.31	2.87	3.31	3.73	3.18	3.15	3.71	4.02
K <sub>2</sub> O	0.98	0.88	0.96	1.03	1.52	0.66	0.75	0.48	0.62	0.89	0.98	1.29	0.91	1.02
P <sub>2</sub> O <sub>5</sub>	0.03	0.01	0.04	0.01	0.20	0.03	0.09	0.07	0.09	0.06	0.08	0.08	0.10	0.12
Total	100.02	99.01	101.13	98.86	96.82	97.27	98.23	99.70	99.27	98.35	99.39	100.35	100.40	99.90
Mg														
Mg + Fe	0.594	0.653	0.550	0.634	0.433	0.558	0.493	0.535	0.641	0.462	0.459	0.502	0.493	0.463
Trace Elements (ppm)														
Ni	76	76	61	97	11	74	29	56	60	24	23	29	26	14
Cr	212	216	185	290	11	167	33	153	62	13	15	23	38	16
Sr	105	103	103	104	169	110	139	119	125	125	108	130	135	134
Zr	33	30	32	38	101	49	62	37	53	68	53	49	59	69
Y	3	2	8	6	31	15	24	15	29	27	23	19	26	32
Nb	2	3	5	7	7	4	5	4	6	6	2	3	4	5
CIPW Norms <sup>c</sup>														
Q	4.31	0.00	10.86	0.82	0.00	6.25	0.91	4.49	0.00	0.52	8.56	8.06	3.54	2.19
Or	5.78	5.19	5.67	6.08	8.97	3.90	4.43	2.83	3.66	5.25	5.78	7.61	5.37	6.02
Ab	21.98	22.24	20.97	20.88	44.24	22.41	27.99	24.27	27.99	31.54	26.89	26.63	31.37	33.99
An	26.90	24.87	24.77	26.51	12.21	25.16	21.12	25.52	22.32	18.55	18.57	18.33	19.67	16.86
Ne	0.00	0.00	0.00	0.00	2.15	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Di	8.54	7.59	9.81	5.42	7.02	6.65	4.49	11.54	7.96	4.63	7.03	7.92	8.59	7.95
En	4.83	4.65	5.21	3.24	3.19	3.49	2.22	6.09	4.89	2.22	3.29	3.96	4.26	3.77
Fs	3.34	2.51	4.29	1.90	3.77	2.96	2.18	5.10	2.61	2.35	3.66	3.78	4.15	4.07
Hy	12.39	15.19	8.89	19.15	0.00	12.10	14.52	8.45	8.47	13.07	9.55	9.90	9.29	9.26
Fs	8.58	8.19	7.32	11.22	0.00	10.26	14.22	7.07	4.52	13.86	10.62	9.45	9.06	9.99
Ol	0.00	3.10	0.00	0.00	4.30	0.00	0.00	0.00	7.53	0.00	0.00	0.00	0.00	0.00
Fa	0.00	1.84	0.00	0.00	5.61	0.00	0.00	0.00	4.44	0.00	0.00	0.00	0.00	0.00
Mt	1.72	1.90	1.69	1.90	1.79	1.95	2.53	1.84	1.98	2.53	2.20	2.01	2.04	2.20
Il	0.61	0.69	0.59	0.67	2.15	0.99	2.10	1.31	1.58	2.30	1.85	1.41	1.73	2.15
Ap	0.07	0.02	0.09	0.02	0.46	0.07	0.21	0.16	0.21	0.14	0.19	0.19	0.23	0.28

<sup>a</sup> Compiled from data of Wood et al. (this volume).<sup>b</sup> Number of analyses in parentheses if more than one.<sup>c</sup> Calculated assuming  $\text{Fe}^{2+}/(\text{Fe}^{2+} + \text{Fe}^{3+}) = 0.86$ .

western Pacific (Natland, crystal morphologies chapter, this volume). Olivine boninites contain, in addition to these minerals, phenocrysts of olivine and clinopyroxene, and fairly abundant magnesiochromite. The highly magnesian pyroxenes as well as the groundmass crystallization sequence allow these rocks to be clearly distinguished from both basalts and andesites (in which plagioclase crystallizes before augite) whether or not they are glassy (Natland, crystal morphologies chapter, this volume).

The two boninite sequences in Hole 458 are designated Types A<sub>1</sub> and A<sub>2</sub> in Figure 7A and Table 2. Chemically, the rocks have the SiO<sub>2</sub> range of typical mafic andesites (52–59%) but high MgO (6–9%), Ni (100–200 ppm), and Cr (200–400 ppm). They have very low TiO<sub>2</sub> (0.3–0.5%), Zr (25–50 ppm), and Y (3–15 ppm). Some chemical variation within Chemical Type A<sub>1</sub> (columns 2–4, Table 2) suggests the presence of subunits, but their boundaries appear to have been obscured by alteration. Besides boninites, there are nine tholeiitic basalt types and subtypes in the two holes (B<sub>1</sub> and B<sub>2</sub> in Hole 458; B<sub>1A</sub>, B<sub>1B</sub>, and B<sub>2A–E</sub> in Hole 459B). Apart from Type B<sub>1A</sub>, which has some chemical and mineralogical features suggesting that it is transitional to boninite compositions (cf. Natland, this volume, and discussion below), these various basalts have the geochemical features of island arc tholeiites (Wood et al., this volume; Meijer

et al., this volume; Sharaskin, this volume), including TiO<sub>2</sub> and Zr lower than in abyssal tholeiites (but higher than in boninites; Table 2), and low Ni and Cr. Several of them are distinctly iron-enriched. They also have generally depleted rare earth element abundances, but moderate abundances of mobile elements such as K, Rb, Ba, and Th (Wood et al., this volume; Hickey and Frey, this volume; Bougault et al., this volume). The two major chemical types, boninite and arc tholeiites, alternate in Hole 458 (Fig. 7A).

Magnetic units shown in Figure 7 are defined on the basis of one or more measurements of magnetic inclination which prove to be nearly identical in stratigraphic order, based on the data of Bleil (this volume). Sequential measurements no different from those that might be expected from the precision of the technique served to define magnetically coherent intervals in the core. These are designated by the average inclinations shown on Figure 7. In the case of both magnetic and chemical units, unit boundaries are assumed to be at the base or top of a core in which only one chemical and magnetic measurement apiece was made. This was the usual case for magnetic measurements, since recovery of oriented specimens was poor for most cores. In some cores, however, unit boundaries are within them. In these cases, the exact boundaries may have been impossible to determine, owing to the difficulty of determining chemistry on ad-



jacent rock pieces (the sampling for shore-based X-ray fluorescence was done onboard ship) and the lack of oriented samples precisely across a magnetic transition. For these, unit boundaries are taken to be any obvious lithologic or petrographic transition in the core between the measured samples.

Chemical units are presumed to represent individual eruptive events. In a number of instances, chemical and magnetic boundaries coincide, implying that the lavas of those particular eruptive events "sampled" the earth's magnetic field at specific points in its history of secular variation and polarity reversals. Given the low latitude of the sites, and a latitude lower still during the Eocene (Bleil, this volume), secular variation totaling about 30° could give rise to apparent reversals, although generally low inclinations would prevail. This could explain the numerous reversals in basalts of Hole 459B. The reversal in Hole 458 between Cores 40 and 41, however, has an average difference in inclinations above and below (weighing each magnetic unit equally) of 37°, in excess of the secular variation. It also occurs across a transition between major chemical types (A<sub>1</sub> and B<sub>1</sub> in Fig. 7). Therefore the eruptive pile could have built up across the period of a reversal in the Earth's magnetic field. Assuming that the reversal itself took about 2000 years to occur (Stacey, 1977), the total time over which the eruptions occurred might have been of the order of  $0.5-1 \times 10^4$  y.

Changes in inclination and apparent reversals could also have been produced by faulting. In both holes, there are magnetic unit boundaries within clearly homogeneous chemical units. Some of these boundaries coincide with zones of intense fracturing in the rocks. Because these are almost certainly fault zones, the magnetic inclinations in part reflect rotation of fault blocks. However, the typical magnitude of such rotations is less than the average 37° difference between the intervals of different polarity in Hole 458 basement. Since that major transition coincides with a chemical boundary as well, it seems likely that a reversal is indeed recorded in basement in Hole 458.

This has important consequences for the magmatic history of the fore-arc region. Both the major lava types that occur in Hole 458 (boninite and arc tholeiite) occur above and below the major magnetic reversal. None of the chemical types is identical to any other (Table 2); hence faulting has not repeated any chemical unit. Consequently, it appears that both magma types erupted alternately throughout a period sufficiently long for a reversal in the Earth's magnetic field to occur. The two chemical types do not belong to any sort of evolutionary sequence (as for example alkalic basalt cappings on tholeiitic oceanic volcanic edifices) but rather represent two distinct compositions supplied to the fore-arc region at much the same time.

This is not just a feature of Hole 458. Chemical data listed in Table 2 and pyroxene compositional data summarized in Natland (crystal morphologies chapter, this volume) show that Chemical Subtype B<sub>1A</sub> of Hole 459B, although petrographically a basalt, resembles boninite Type A<sub>2</sub> of Hole 458. Both boninitic and basaltic lava

fragments were obtained in talus deposits drilled beneath Eocene sediments (Ellis, this volume) deep in the Mariana Trench at Sites 460 and 461 (Sharaskin, this volume; Meijer et al., this volume). Lavas from these holes have arc tholeiite compositions (Wood et al., this volume). Boninites and "marianites" (which are boninites with abundant clinoenstatite) were also obtained in a dredge haul in the Mariana Trench near Guam (Dietrich et al., 1978; Sharaskin et al., 1980). These were associated with an "ophiolitic" assemblage of gabbros, norites, graywackes, and ultramafic rocks. A "low-TiO<sub>2</sub> basalt" from this dredge (compositionally a boninite) has a K/Ar age of 10.8 m.y. (Beccaluva et al., 1980). Another nearby dredge haul recovered island arc tholeiites (Sharaskin et al., 1980). Boninites and island arc tholeiites have been recovered in several other dredge hauls in the trench and form several fore-arc basement highs similar to that near Site 458 (S. Bloomer and J. W. Hawkins, personal communication). Meijer (1980) has pointed out that boninites series lavas occur in the Miocene Umatuk Formation on the island of Guam. These overlie Eocene basalts and andesites and are associated with Miocene calc-alkalic lavas. Stark (1963) describes the boninitic lavas as basalts but notes that they are "hypersthene"-bearing, on the basis of parallel extinction, with hypersthene occurring in glomerophytic clots with augite. Plagioclase microlites are abundant in groundmass. These are the petrographic features of the more holocrystalline Hole 458 boninite pillows, although the latter contain bronzite phenocrysts (Natland, crystal morphologies chapter, this volume; Bougault et al., this volume; Meijer et al., this volume; Sharaskin, this volume). For comparison, chemical analyses of the Guam boninites are provided in Table 3.

In summary, we conclude that the boninite and arc tholeiite lava series were supplied to the Mariana arc and fore-arc region essentially simultaneously at most

Table 3. Compositions of boninites from Guam compared with a representative Hole 458 boninite.

Umatuk Formation (Guam) Boninite-type Lavas <sup>a</sup>					Sample 458-39-2, 94-96 cm <sup>b</sup>
SiO <sub>2</sub>	50.85	51.63	54.2	55.0	52.5
Al <sub>2</sub> O <sub>3</sub>	13.55	14.10	14.6	13.8	15.7
Fe <sub>2</sub> O <sub>3</sub>	2.24	2.27	5.1	3.8	9.43 <sup>c</sup>
FeO	5.43	5.93	3.4	4.0	
MgO	10.01	9.93	8.0	9.4	9.16
CaO	9.55	9.47	7.9	7.9	8.39
Na <sub>2</sub> O	1.58	2.21	2.4	2.3	2.63
K <sub>2</sub> O	0.16	0.52	1.4	0.88	0.86
TiO <sub>2</sub>	0.34	0.43	0.36	0.34	0.36
P <sub>2</sub> O <sub>5</sub>	0.03	0.05	0.08	0.07	0.01
MnO	0.14	0.15	0.12	0.11	—
H <sub>2</sub> O <sup>+</sup>	1.91	1.42			—
H <sub>2</sub> O <sup>-</sup>	4.13	1.84	2.9	2.6	—
Total	99.92	99.95	100.46	100.20	99.20
Ni	300	200			99
Cr	900	700			217
Zr	10	10			35
Y	30	30			<1
Sr	200	300			102
Ba	20	70			20

<sup>a</sup> Data from Tracy and Stark (1963).

<sup>b</sup> Data from Wood et al. (this volume).

<sup>c</sup> Total iron as Fe<sub>2</sub>O<sub>3</sub>.

locations in the Eocene, and again on the island of Guam during the Miocene. There is no indication in the drill sites or in any of the dredge hauls that genuine ocean crust exists either beneath the fore-arc sediments or at very great depths in the Mariana Trench. Such crust would have to be at least Early Cretaceous, and probably Jurassic, in age (Hussong and Fryer, this volume) and be overlain by some thickness of pelagic sediments. The fore-arc region is therefore not buttressed by major fault slices derived from the converging Pacific lithospheric plate.

#### Alteration in Fore-arc Basement

Many of the lavas in Holes 458 and 459B are extensively altered, especially along zones of intense fracturing of the rocks. A variety of clays and zeolites formed, and—mainly in Hole 459B—an unusual low-Al, high-Fe form of palygorskite (Natland and Mahoney, this volume) occurring exclusively in veins. The alteration was in two stages in both holes: (1) an early oxidative, possibly hydrothermal stage during which dioctahedral smectites, celadonite, palygorskite, calcite, and iron hydroxides formed and (2) a later, less oxidative stage in which di- and/or trioctahedral smectites and phillipsite formed. The latter occurred especially along zones of intense fracture in the lavas, which probably formed as a result of faulting of the fore-arc region. Profiler records indicate extensive normal faulting of the fore-arc, especially near Site 459 (Mrozowski and Hayes, 1980), where the sense of faulting is predominantly toward the trench. The faulting appears to be occurring even today, since Recent sediments have been affected by it (Hussong and Fryer, this volume). Takigami and Ozima (this volume) have obtained an apparently reliable age for the second-stage alteration in Hole 458 by  $^{40}\text{Ar}/^{39}\text{Ar}$  incremental heating techniques. This age is about 10 m.y. younger than the age of the oldest (early Oligocene) sediments at the site.

The palygorskite in Hole 459B is its first reported occurrence in igneous basement. Elsewhere in the oceans, it has been reported only in pelagic or volcanoclastic sediments. Natland and Mahoney (this volume) speculate that it formed under oxidative hydrothermal conditions and that the second stage of alteration was caused by fluids circulating in the basement rocks whose compositions may have been different from those of fluids normally found in ocean crust. An important component of the fluids may have been derived from subducting ocean crust and sediments now only a few kilometers beneath the fore-arc region. The fluid compositions may have been additionally modified by having to travel along faults through a thick assemblage of mafic and ultramafic rocks similar to those drilled and dredged in the landward wall of the Mariana Trench.

#### Geochemical Relationships among Lavas of Holes 458 and 459B

Contributors to this volume generally agree that there is a spectrum of forearc parental types produced by the combined effects of different degrees (and possibly depths) of partial melting, and source heterogeneities.

There is also a general consensus that the sources differ from those of abyssal tholeiites, principally by addition of mobile constituents derived from the subducted ocean crust and sediments. Finally, it is apparent that the oxidation state of the magmas influenced fractionation, producing different degrees of iron,  $\text{TiO}_2$ , and  $\text{SiO}_2$  enrichment.

These conclusions can be demonstrated using simple variation diagrams, such as those of Figures 8 and 9 on which the averaged data of Table 2 are plotted. Figure 8A,  $\text{TiO}_2$  versus total iron as  $\text{Fe}_2\text{O}_3$  ( $\text{Fe}_2\text{O}_3^T$ ), shows that the boninites of Hole 458, as well as basalt Type B<sub>1A</sub> of Hole 459B, have systematically lower  $\text{TiO}_2$  and  $\text{Fe}_2\text{O}_3^T$  than the remaining basalts. Moreover, with only one exception, both boninites and basalts have systematically lower  $\text{TiO}_2$  at a given  $\text{Fe}_2\text{O}_3^T$  than abyssal tholeiites from the East Pacific Rise and Galapagos-Costa Rica Rifts showing comparable variation in  $\text{Fe}_2\text{O}_3^T$ . This is one of the criteria by which Jakes and Gill (1970) distinguished island arc tholeiites from mid-ocean-ridge basalts. The single exception on Figure 8A is basalt Type B<sub>1</sub> from Hole 458 (Table 3), which has unusually high  $\text{Na}_2\text{O}$ , low  $\text{MgO}$ , and low  $\text{SiO}_2$ . Its composition is based on a single analysis of rocks probably greatly modified by alteration. Even so, its  $\text{TiO}_2$  content compares only with the lowest  $\text{TiO}_2$  abundances found in abyssal tholeiites.

Three groups of rocks can be distinguished on a plot of  $\text{Fe}_2\text{O}_3^T$  versus  $\text{SiO}_2$  (Fig. 8B): a group of four boninites (here including Hole 459B Subtype B<sub>1A</sub> and the average boninite andesite of Column 3, Table 3); a group of four low- $\text{SiO}_2$  tholeiites, two of which have marked iron enrichment (Type B<sub>2</sub> of Hole 458 and Subtype B<sub>2A</sub> of Hole 459B); and a group of four high- $\text{SiO}_2$  tholeiites (Subtypes B<sub>2B-E</sub> of Hole 459B) which have less iron enrichment. They are also lower in  $\text{TiO}_2$  than the two most iron-enriched low- $\text{SiO}_2$  tholeiites.

The variation diagrams of Figure 9 use Zr on the abscissa as an index of differentiation unaffected either by alteration or by oxidation state of the magmas. On several of these diagrams, the three major groups fall into different fields, and on some they overlap. For reference, the least fractionated arc tholeiite (Subtype B<sub>1B</sub> of Hole 459) and the most iron-enriched from the same hole (Subtype B<sub>2A</sub>) are connected by a solid line. The most fractionated basalt type, with the highest Zr and Y, and the lowest normative  $\text{An}/\text{An} + \text{Ab}$ , is Type B<sub>1</sub> of Hole 458, the altered sample mentioned earlier. Zr, Y, and  $\text{TiO}_2$  should not have been much affected by the alteration. Consequently, this basalt has surprisingly lower  $\text{TiO}_2$  and  $\text{Fe}_2\text{O}_3^T$  than the two iron-enriched basalts, even though it appears to be more fractionated on the basis of trace elements. Either iron enrichment among its precursors was not as extreme, there was a marked late-stage fractionation of titanomagnetite (contrasting dashed arrows on Figs. 8 and 9), the source of its parent in the mantle was less depleted than the source of the other basalts, or its parent was a smaller partial melt.

The high- $\text{SiO}_2$  basalts overlap the iron-rich tholeiites in Y, Zr, and  $\text{TiO}_2$  on Figure 9 variation diagrams but have lower  $\text{Mg}/\text{Mg} + \text{Fe}$  and normative  $\text{An}/\text{An} + \text{Ab}$ .

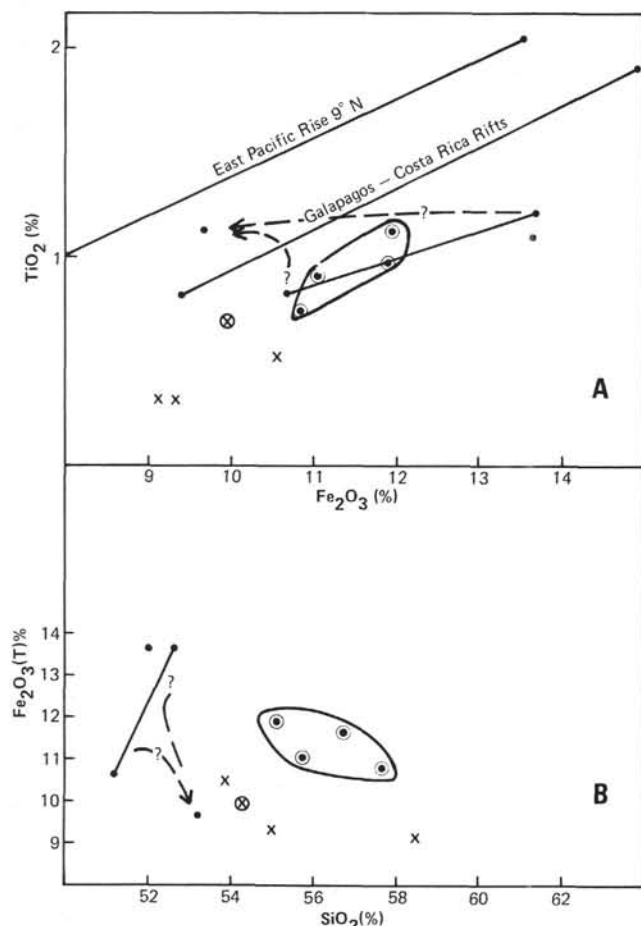


Figure 8. A.  $\text{Fe}_2\text{O}_3(\text{T})$  versus  $\text{TiO}_2$  for basalts and boninites of Holes 458 and 459B, using average data of Table 2. Trend lines for East Pacific Rise and Galapagos-Costa Rica rifts fractionation trends are based on glass compositions in Natland and Melson (1980) and unpublished data. (X = boninites, ⊗ = Subtype B<sub>1A</sub> of Hole 459B, • = low  $\text{SiO}_2$  tholeiites, and ⊙ = high  $\text{SiO}_2$  tholeiites.) B.  $\text{SiO}_2$  versus  $\text{Fe}_2\text{O}_3(\text{T})$  for average data of Table 2. (Same symbols as A. Solid lines on A and B connect least and most fractionated Hole 459B low  $\text{SiO}_2$  tholeiites. Dashed arrows are possible alternative paths of fractionation to Hole 458 Type B<sub>1</sub> basalt.)

They have 2.5%–6.5% higher  $\text{SiO}_2$ . Within the four high- $\text{SiO}_2$  basalts, however,  $\text{SiO}_2$  decreases with increasing Zr, Y,  $\text{TiO}_2$ , and  $\text{Fe}_2\text{O}_3^{\text{T}}$ , contrary to what might be expected from a normal fractionation sequence. The normative ratio  $\text{An}/\text{An} + \text{Ab}$ , however, decreases in a way consistent with fractionation.

These are complex variations, unlike the simple trend toward iron enrichment shown by eastern Pacific ferrobasalt suites (cf. Clague and Bunch, 1976; Natland and Melson, 1980). If all the tholeiitic basalts are related to a single parent, then some of the variation in  $\text{TiO}_2$ ,  $\text{Fe}_2\text{O}_3^{\text{T}}$ , and  $\text{SiO}_2$  might be related to variable fractionation of titanomagnetite, together with plagioclase and clinopyroxene. This would simultaneously suppress enrichment in  $\text{Fe}_2\text{O}_3^{\text{T}}$  and  $\text{TiO}_2$  and promote  $\text{SiO}_2$  enrichment (Osborne, 1959, 1962, 1979). The parent, however, would have to be far more magnesian and less siliceous than the least-fractionated basalt, Subtype B<sub>1B</sub> of Table 2. Moreover, the extent of titanomagnetite frac-

tionation would not be the same for any of the high- $\text{SiO}_2$  tholeiites, since the least evolved of those in terms of Y and Zr (Subtype B<sub>2C</sub>) has the highest  $\text{SiO}_2$ .

There is little reason, however, to suppose that a common parent was involved. Instead, the low- $\text{SiO}_2$  tholeiites have systematically higher Zr at a given Ni contents (Fig. 9H) than the high- $\text{SiO}_2$  tholeiites; hence one might suppose that the parent to the latter represents a greater degree of melting of a homogeneous source than the parent of the low- $\text{SiO}_2$  tholeiites. This is based on an argument by Bougault et al., (1978) that Ni is insensitive, whereas elements such as Zr are quite sensitive, to variations in the degree of melting in an olivine-rich peridotitic source. In terms of most of the oxides and trace elements considered in Figures 8 and 9, we can even say that such a parent would have approached boninite compositions and that the high- $\text{SiO}_2$  tholeiites would be iron-enriched with respect to such a parent.

We therefore propose that the mantle source of the Mariana fore-arc region produced a spectrum of parental compositions ranging from olivine boninite to olivine tholeiite and that different oxidation states in the tholeiite magmas produced variable  $\text{SiO}_2$  enrichment and  $\text{TiO}_2$  and iron depletions. There is some petrographic evidence for this. Titanomagnetite is particularly abundant in Hole 459 tholeiites (indeed, it is so abundant that it produces intensities of magnetization 3–5 times higher than in even more iron-enriched abyssal tholeiites; Bleil, this volume). Some of the basalts contain segregation vesicles in which titanomagnetite is the only well-crystallized phase, indicating high oxygen fugacities in the vesicles (Natland, crystal morphologies chapter, this volume). Exactly at what stages titanomagnetite entered fractionation sequences is not clear. This may have depended on concentrations of volatiles in the mantle sources. What is clear is that its impact was quite variable, possibly reflecting conditions in magma chambers and conduits that held these magmas. Some of the chemical contrasts among the basalt types could also have been caused by mixing variably fractionated and oxidized magma batches. Whatever the case, it now seems that in consideration of a suite of rocks, variable iron,  $\text{TiO}_2$ , and  $\text{SiO}_2$  enrichments may characterize many island arc tholeiites and be useful as discriminants between these and ocean floor basalts. A given arc tholeiite may not be particularly iron-enriched, or indeed much fractionated, but it still can be distinguished from abyssal tholeiites by having lower  $\text{TiO}_2$  at a given total iron abundance (Fig. 8A).

Both the basalts and the boninites of Holes 458 and 459B have fractionated compositions. No compositions potentially representative of parental compositions were recovered. This has been explicitly determined experimentally for the Hole 458 boninites by Kushiro (this volume), who inferred that they did not melt in equilibrium with a hydrous plagioclase lherzolite mantle but instead seem to be derived from an olivine boninite parent which did. This is also indicated by the absence of both clinoenstatite and olivine in Hole 458 boninites and their somewhat low MgO contents.



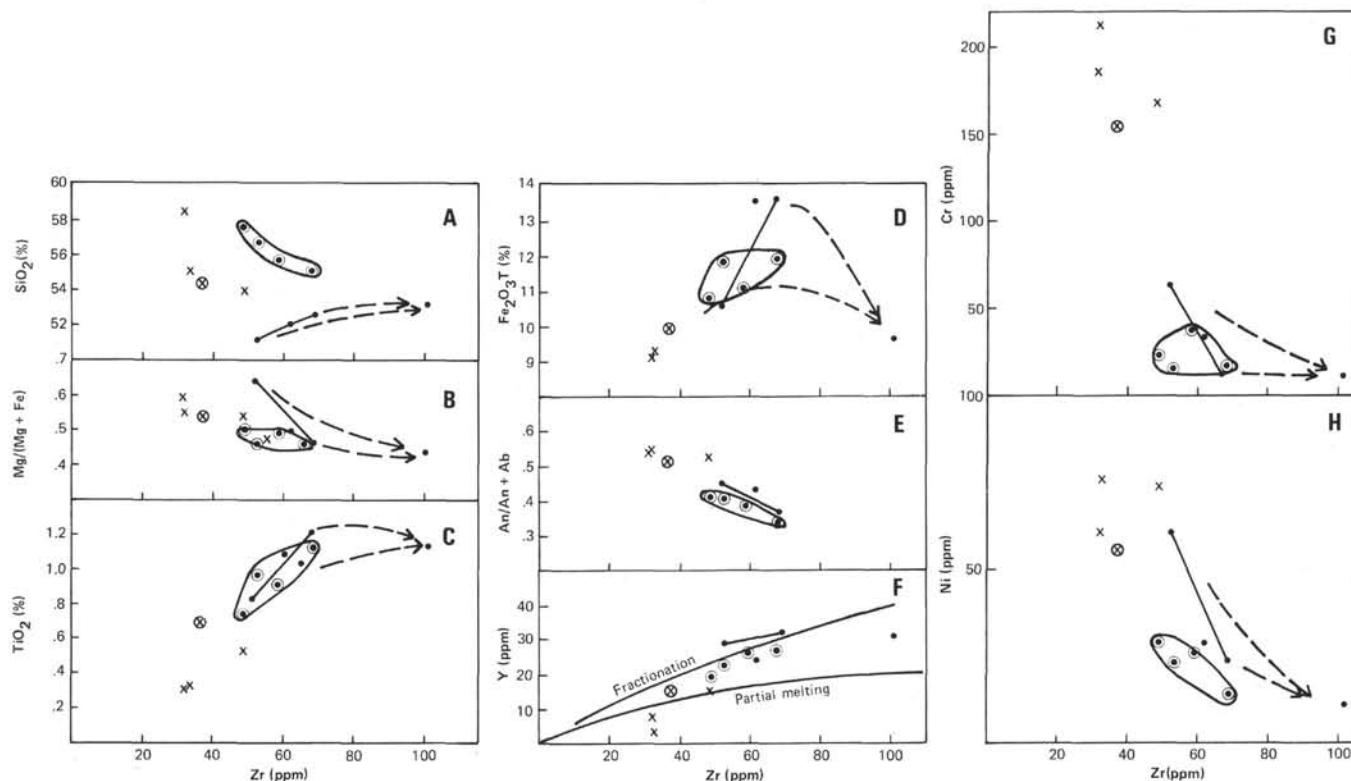


Figure 9. Zr variation diagrams for average chemical types and subtypes of Holes 458 and 459B, from Table 2. (Symbols as in Figure 8A). A.  $\text{SiO}_2$ -Zr; B.  $\text{Mg}/(\text{Mg} + \text{Fe})$ -Zr; C.  $\text{TiO}_2$ -Zr; D.  $\text{Fe}_2\text{O}_3(\text{T})$ -Zr; E. Normative  $\text{An}/(\text{An} + \text{Ab})$ -Zr; F. Y-Zr with fractionation and partial melting trends of MORB suites from Tarney et al. (1979); G. Cr-Zr; H. Ni-Zr.

### Contrasts to MORB and the Compositions of Fore-arc Mantle Sources

Data contributed to this volume and to DSDP Volume 59 (Kroenke, Scott, et al., 1980) add to the growing body of information on the distinctive geochemistry of island arc lavas. As summarized from Wood et al. (this volume), Bougault et al. (this volume), and Hickey and Frey (this volume), the island arc tholeiites of Holes 458 and 459B differ from MORB in the following ways: (1) They have lower Ti and Zr and higher Ti/Zr at a given MgO content despite equivalent or greater iron enrichment; (2) Overall, they have lower rare-earth element abundances despite being generally more fractionated. However, they have similar or even somewhat greater light rare-earth element depletion; (3) Although they have lower Zr, Ti, and Hf than MORB, the abundances of these elements are not as low relative to the depleted rare-earth element abundances as in MORB; (4) Ta and Nb are even more depleted than rare-earth elements when compared with MORB; (5) Highly incompatible elements such as Cs, K, Rb, Ba, Pb, U, and Th are enriched in the arc tholeiites compared with MORB and their own depleted rare-earth element abundances; (6) The arc tholeiites have higher  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7036–0.7038) than MORB (Armstrong and Nixon, 1980), although some of this may have resulted from alteration.

Despite having erupted apparently contemporaneously with the arc tholeiites, the Hole 458 boninites cannot have been derived from a similar mantle source,

although they, too, differ significantly from MORB. Their very low abundances of Ti, Zr, rare-earth elements, and other highly incompatible elements and their high MgO, Ni, and Cr indicate that their presumed olivine boninite parent was a high partial melt of the mantle. Wood et al. (this volume), Meijer et al. (this volume), and Kushiro (this volume) argue from various standpoints that melting occurred under hydrous conditions and that the source was refractory (had experienced a previous melting episode of basalt extraction).

Normally, one would expect that combining high partial melting and a refractory source, would result in considerable depletions in highly incompatible elements; not so with these lavas. They are even more enriched, relative to their rare-earth element abundances, in large ion lithophile elements such as K, Rb, U, and Th than the associated arc tholeiites, or than MORB. In this respect, they display the characteristics of calc-alkalic suites. They show an even more marked enrichment of Zr relative to Ti than the arc tholeiites (although overall abundances of both are quite low). However, they do not show the depletion in Ta and Nb relative to rare-earth elements of the arc tholeiites.

When the spectrum of magmatic processes that have acted on these volcanic rocks is put in perspective, it is evident that although fractionation and even melting processes have been important, the distinctive features of these rocks relate to their source compositions. Petrogenetic models proposed by contributors to this volume quite naturally have this as their major concern.



There is, for example, agreement among Wood et al. (this volume), Hickey and Frey (this volume), and Bougault et al. (this volume) that the mantle sources have been enriched in large ion lithophile elements, probably by fluids derived from the subducting ocean crust and its sediments (e.g., Hawkesworth et al., 1979; Saunders et al., 1980; Kay, 1980). Wood et al. (this volume) explain the refractory behavior of Ti, Ta, and Nb by arguing that a Ti-rich phase such as rutile or sphene is stable under hydrous conditions and retains these elements during the partial melting of a hydrous mantle source (Hellman and Green, 1979). The lack of Ta and Nb depletions in the high-MgO boninite andesites, they suggest, reflects the exhaustion of such a phase or phases during extreme partial melting. The unusual enrichment in Zr and Hf relative to Ti, Y, and the heavy rare-earth elements is more difficult to explain by crystal-melt relationships and would require retention of Ti and the heavy rare earths in garnet during melting (Wood et al., this volume) or replenishment of the mantle source with incompatible elements such as Zr and Hf in addition to K, Rb, U, Th, etc.

### Models for Fore-arc Magmatism

Reconciling these geochemical features with plausible tectonic-magmatic sequences in the Mariana arc system has not produced unanimity of opinion. Most petrologic contributors to this volume have seized on the Eocene-Oligocene age of the boninite series lavas in the Bonin Islands and at Site 458 as an indication that these unusual lavas are somehow restricted to eruption during the earliest stages of arc volcanism only. Meijer et al. (this volume) and Meijer (1980), for example, suggest that at the onset of subduction, cold hydrous ocean crust is suddenly plunged into hot mantle. The consequent release of water from the subducted crust into the hot mantle suppresses the liquidus sufficiently to initiate melting, and a spectrum of magma types is produced as the zone of melting rises from the vicinity of the subduction zone to shallower levels in the mantle wedge above it. Boninites are produced when melting reaches as shallow as 40 to 60 km below the seafloor. As subduction continues, the temperature of the mantle around the subducted lithosphere is lowered, and because the zone of melting cannot rise to such shallow depths, boninites cease to erupt. This hypothesis, however, does not take into account either the early Miocene boninites on Guam or late Miocene (10.8 Ma) boninite lavas dredged near Guam. For the latter, Beccaluva et al. (1980) propose that the thermal conditions necessary to produce boninites were reestablished by a rise in isotherms at the initiation of Mariana Trough back-arc basin spreading. This could have caused additional melting of shallow, but nevertheless refractory, subarc mantle. Hickey and Frey (this volume) and Wood et al. (this volume) have discussed models in which the mantle above the ocean crust becomes progressively enriched in the highly incompatible elements as subduction proceeds. Hickey and Frey relate this specifically to the sequence of volcanism recorded in Hole 458, in which the relatively more enriched boninites rest stratigraphically above the

arc tholeiites. Meijer (1980) made the similar point that boninite series lavas seem to occur later in volcanic successions on Guam and the Bonin Islands as well as in Hole 458. In our view, the evidence for later eruption of boninites is circumstantial and based on extremely restricted outcroppings or drill holes. In Hole 458, the two magma series clearly alternate. Matthey et al. (1980) and Wood et al. (this volume) make the more general point that the Mariana arc as a whole has evolved to more calc-alkalic compositions, especially on the West Mariana Ridge and the currently active arc, and that this entails steady enrichment in the mantle of the mobile incompatible elements. They mention the similarities between boninite series and calc-alkalic series lavas in these elements. Hickey and Frey (this volume) demonstrate that boninites show quite variable enrichments in light rare-earth element abundances. This suggests to them that reenrichment of depleted mantle sources by the highly mobile incompatible elements can result in a continuum of depleted to enriched boninite series lavas.

Meijer et al. (this volume), Wood et al. (this volume), and Hickey and Frey (this volume) all favor an origin for arc tholeiites in the mantle above the subducting oceanic lithosphere. Of these authors, only Wood et al. (this volume) propose that the depleted source of the boninites was originally the uppermost mantle of the subducting lithosphere from which abyssal tholeiites had been extracted. Upon deep descent into the mantle during subduction, this refractory peridotite would become lighter than surrounding mantle from which basaltic melts had not been extracted, probably because it would contain comparatively less of the dense iron-rich phase, garnet. It would then ascend diapirically into the mantle above the subducting lithosphere, there to become variably reenriched in mobile elements also derived from the subducting crust.

This proposal is similar to that of Kay (1980), who favored derivation of the entire suite of island arc magmas from rising depleted peridotite diapirs originating from within the subducting lithosphere, based on the low abundances of Ti, Nb, Zr, and heavy rare-earth elements in all of them (Green, 1976; Sun and Nesbitt, 1978). Variable contamination with subducted sediments and seawater again would cause the enrichments in highly mobile incompatible elements. The magmas could also be modified to a greater or lesser extent by crystal fractionation prior to eruption.

On geochemical grounds, however, it is still impossible to distinguish between models favoring subduction zone melting (e.g., Ewart et al., 1977), shallower melting of depleted diapirs rising from the subducting lithosphere (Kay, 1980), or melting of the mantle above the subduction zone after it has been hydrated and variably enriched in highly mobile incompatible elements (Ringwood, 1974; Thorpe et al., 1976; Hawkesworth et al., 1979; Saunders et al., 1980). All of these models require ways of simultaneously depleting melts in Ti, Zr, Hf, Ta, Nb, and, sometimes, P, and enriching in K, Rb, Ba, Th, U, Cs, and radiogenic strontium. The depletions are accomplished either by a two-stage melting process, in which a previous extraction of basalt has occurred, or

by stabilization under hydrous conditions of minor phases such as rutile, sphene, and apatite, which retain the relatively depleted elements during melting. Saunders et al. (1980) argue that the latter could occur anywhere near the hydrous part of the subducting crust, leaving only the mobile incompatible elements to escape during dehydration. Once this has occurred, however, it would seem that melting could occur either in the subduction zone or in rising diapirs and produce the same geochemical anomalies, provided adequate enrichment or re-enrichment in the mobile elements also occurs.

Experimental evidence for the actual depth of segregation of island arc parental materials from the mantle favors shallow depths rather than the great depths of typical subduction zones. Experimental data indicate that boninites could have melted in equilibrium with hydrous peridotite at depths corresponding to about 10–20 kbar (e.g., Kushiro, 1972, 1974, and this volume; Green, 1973, 1976; Nicholls and Ringwood, 1973; Nicholls, 1974; Mysen et al., 1974; Mysen and Boettcher, 1975a, 1975b). In their review of the chemical characteristics of island arc basalts, Perfit et al. (1980) conclude that the major oxide differences between arc basalts and MORB are not a consequence of differences at the source but appear to reflect different crystallization histories: Notably, arc basalts have less initial plagioclase fractionation and more clinopyroxene fractionation. This, these authors argue, is a consequence either of higher  $p(\text{H}_2\text{O})$  or greater depths of crystallization. We have argued that variable oxidation states of arc basaltic magmas cause different enrichments in total iron,  $\text{TiO}_2$  and  $\text{SiO}_2$ . This might be related to a more hydrous mantle source than that of MORB, but this does not imply a greater depth of melting. There is no important compositional feature of these basalts to indicate depths of melting significantly greater than those of ocean floor basalts. The latter have been interpreted to originate at depths of about 30 km (9–10 kbar; e.g., Kushiro, 1973; Fujii et al., 1978; Bender et al., 1978; Presnall et al., 1979) or perhaps twice that much (e.g., Duncan and Green, 1980; Stolper, 1980; Jaques and Green, 1980). Similar depths of origin seem preferred for boninites, as already discussed. These conclusions are consistent either with melting of peridotite diapirs rising from the subducting lithosphere or with the direct melting of mantle above the subduction complex itself, following its hydration and enrichment in mobile incompatible elements.

The two magma types, arc tholeiites and boninites, can be viewed as complements in a direct melting model, with boninites derived from the refractory mantle produced by extraction of the former. The situation would be analogous to that of the FAMOUS area on the Mid-Atlantic Ridge, where continued melting of a mantle from which basalts have been extracted has been invoked to explain the interbedding or close spatial proximity of sparsely plagioclase phyric and strongly olivine phyric basalts (Langmuir et al., 1977; Natland, 1978; Duncan and Green, 1980). The differences would be (1) that the mantle beneath the Mariana fore-arc was more

hydrous and (2) that variable external enrichment in mobile elements was occurring. Based on Holes 458 and 459B, it is our belief that olivine boninites and olivine-bearing parents to arc tholeiites are end components of a continuum of parental compositions supplied to the Mariana fore-arc region during late Eocene–early Oligocene times. Whether intermediate magma types may have been produced by hybridization between these end components is uncertain. But there is no reason why many of them could not have been derived directly from variably refractory and hydrated sources in the mantle. The existence of such a continuum implies that the end component magma types were simultaneously available, as argued earlier on stratigraphic grounds, and probably means that their sources were not too greatly separated spatially in the mantle. In other words, a deep source for arc tholeiites, and a shallow source for boninites (e.g., Meijer, 1980), seems unlikely.

### Boninite Constraints on Arc-Trench Geometry

The eruptive sources of lavas to Sites 458 and 459 were probably not far away, certainly not as far away as the present Mariana arc or even the ancestral Mariana arc which may have been on the northern extension of the ancestral Guam–Saipan arc (Fig. 1). This would have been over 100 km to the west. If we assume that melting occurred more or less directly beneath the two sites, then great changes in the geometry of subduction must have occurred since the Eocene. The present depth to the top of the subducted oceanic lithosphere beneath Site 459 is no more than 20–30 km (Hussong and Fryer, this volume), shallower than the experimentally estimated depth of segregation of boninites. If, as argued by Wood et al. (this volume), boninites and possibly island arc tholeiites arose from diapirs of hydrous peridotite emanating from the subducted lithosphere that ascended through a denser garnet-bearing mantle, then the subduction zone was originally 100 km or more beneath Site 458 and 459. Even if melting was confined to shallower mantle never invaded by such diapirs, the depth to the subduction zone would still have had to be considerably more than it is now. We believe that this is consistent with tectonic removal, since the Eocene, of portions of the eastern edge of the fore-arc region by processes related to subduction. There is evidence from seismic profiler records for present-day downfaulting of the edge of the fore-arc into the Mariana Trench (Mrozowski and Hays, 1980; Hussong and Fryer, this volume), and it is evident that arc-related igneous and metamorphic rocks crop out in the trench wall (e.g., Dietrich et al., 1978; Beccaluva et al., 1980; Sharaskin et al., 1980; Meijer et al., this volume; Sharaskin, this volume; Wood et al., this volume; Sites 460 and 461 chapters, this volume; S. Bloomer, personal communication). The result of this tectonic process has been to shift the zone of subduction westward relative to Sites 458 and 459, and probably even farther to the west at the northern and southern ends of the Mariana arc, where the fore-arc region is narrower than near 18°N.



## Mariana Fore-arc Ophiolite

The Mariana fore-arc is a plateau-like region between the high subaerial volcanoes of the Mariana arc and the abrupt chasm of the Mariana Trench. Beneath the sediment drape which thickens toward the arc, fore-arc basement averages between 4 km and 5 km below sea level, with local elevations as shallow as 1.5 km. On the basis of Leg 60 drilling, the top of this basement appears everywhere to have been produced by Eocene-late Oligocene arc volcanism. The drilling also demonstrated that this region has always been in fairly deep water and, except possibly locally, has never been uplifted. Within and near the Trench, in fact, subsidence has clearly occurred, as evidenced by the Trench Sites 460 and 461, which have subsided below the carbonate compensation depth since the late Oligocene (see Sites 460 and 461 chapters, this volume). Normal faults are abundant near the Trench (Mrozowski and Hayes, 1980; Mrozowski et al., this volume); some even affect Recent sediments (Hussong and Fryer, this volume).

Seismic refraction studies show that the crust of the fore-arc is somewhat thicker and has lower velocities than either normal ocean crust or back-arc basin crust. However, it is not grossly different from either. Velocities approach 5.6–6.1 km/s uniformly at depths of 4.5 km below the seafloor throughout the fore-arc region (Hussong and Fryer, this volume). But "on the basis of seismic velocities only, it is impossible to determine whether the uppermost layer 3 material beneath the fore-arc is composed of arc volcanics or is old back-arc basin or oceanic crust that has been buried by, and now serves as the foundation for, more recent arc volcanics" (LaTraille and Hussong, 1980, p. 215).

The total thickness of the crust increases to the west beneath the Mariana arc (Hussong, this volume, Plate 2, back pocket). The crust of the Palau-Kyushu Ridge is of comparable thickness (Murauchi et al., 1968). There is now, and there was in the Eocene, a thicker crustal root beneath shallow portions of the arc system, where there are, or were, lines of active shallow or subaerial volcanoes.

The structural relationship between the fore-arc lavas and those of the higher Eocene volcanoes of Guam, Saipan, and the Palau-Kyushu Ridge is not clear. But the recovery of gabbros and ultramafic rocks in the wall of the Trench (Dietrich et al., 1978; Beccaluva et al., 1980; Sites 460 and 461 chapters, this volume; S. Bloomer and J. Hawkins, personal communication), in many cases associated with or close to evident outcrops of arc tholeiites and boninites, is strong evidence that fore-arc crust is quite thin and lacks any component of faulted-in ocean crust. The arc volcanics appear directly to overlie gabbros and ultramafic rocks, and it is logical to presume that the gabbros and cumulus ultramafic rocks consolidated in fore-arc rather than in ocean crust magma chambers. Put another way, for the gabbros and ultramafic rocks to have originated in the ocean crust, one would have to suppose the existence of a fault

or faults which removed all vestige of several hundred meters of Jurassic and younger marine sediments and of several kilometers of Layer 2 pillow basalts from the top of the ocean crust section and subducted these without subducting the deeper and denser gabbros and ultramafic rocks. This would probably also have caused uplift of the outer part of the fore-arc region, whereas subsidence has occurred. The arc lavas would now be in fault or depositional contact with the plutonic rocks.

We have summarized evidence that the arc tholeiites in particular, but also the boninites, of Holes 458 and 459B are quite fractionated lavas. A complementary suite of cumulus gabbros must exist somewhere in the fore-arc crust. Earlier we noted that calc-alkalic gabbros were dredged from the Mariana Trench near Guam (Dietrich et al., 1978). It is possible that Seismic Layer 3 of the fore-arc is made up almost entirely of cumulus gabbros involved in the fractionation of arc tholeiites, boninites, and even calc-alkalic lavas.

In short, we propose that the Mariana fore-arc region is an *in situ* ophiolite succession with appropriately thin crust, a reasonably typical velocity structure (probably modified by serpentinization and the extensive faulting that has occurred), and the necessary upper pillowed extrusive sequence overlying appropriate plutonic rocks. It was produced during the early stages of arc volcanism and is not a fragment of ocean crust.

Exactly how this crust was produced must remain a matter for speculation. Cameron et al. (1979, 1980) have pointed to the occurrence of boninites in some ophiolites including Troodos and, because of the boninites dredged and drilled from the Mariana Trench and fore-arc, proposed that they formed in fore-arc settings. Troodos is one ophiolite that has a particularly well-developed sheeted dike complex (Gass, 1963; Moores and Vine, 1971; Kidd, 1977; Coleman, 1977), which is normally taken to indicate an origin by seafloor spreading. Miyashiro (1973) argued for an island arc rather than ocean crust origin for Troodos on the basis of geochemistry, a suggestion that was not well received primarily because Troodos simply does not look structurally like an island arc but instead resembles ocean crust, or at least some type of back-arc basin. The discovery of boninites in the Troodos complex pillow lavas implies that Miyashiro (1973) was right, as Cameron et al. (1980) have argued, but what now should be made of Troodos's ocean crust type of structure?

We suggest that an appropriate equivalent of Troodos in an island arc setting is the Mariana fore-arc region and that at the earliest stages of arc volcanism, a type of seafloor spreading may have occurred before volcanism focused along the line of the arc. Such an early spreading episode is not unreasonable in a tectonic setting that has twice subsequently experienced arc rifting and formation of back-arc spreading basins. Whether or not true sea-floor spreading produced the Mariana fore-arc region remains to be seen in the light of more detailed geophysical experiments, careful dredging in the Trench, and perhaps drilling. Regardless, the occurrence of bo-

ninites in certain ophiolites is good evidence that spreading has occurred in several ancient arc systems, probably early in their history.

## PART II. STRUCTURAL AND MAGMATIC EVOLUTION OF THE MARIANA ARC SYSTEM

### Trapped Ocean Crust, Inactive Arcs, and Back-arc Basins in the Philippine Sea

Here we review some of the more important results of Legs 58 and 59 and integrate them with the Leg 60 data summarized in Part I to provide an overall view of the evolution of the Mariana arc system. The two earlier legs provided valuable information on the West Philippine Sea, the Palau-Kyushu Ridge, and the Parece Vela Basin, information which is important in understanding the early stages of arc development. The location of sites drilled during Legs 58 and 59 are shown on Figure 1. Details of the petrology of igneous rocks recovered during these legs are in numerous chapters in Klein, Kobayashi, et al. (1980) and in Kroenke, Scott, et al. (1980).

#### West Philippine Basin

Hilde et al. (1977) have proposed that the West Philippine Basin (Fig. 1), lying between the Ryukyu Arc at the Asian plate boundary to the west and the Palau-Kyushu Ridge to the east, is "trapped" in origin, being part of normal Pacific ocean floor which became separated from the main ocean as a result of initiation of subduction along the line of the present Palau-Kyushu Ridge. They suggested that the eastern boundary of the West Philippine Sea may initially have been an old N-S transform fault which changed to a convergent plate boundary consequent upon a change in direction of Pacific plate motion from northerly to westerly some 45 Ma (cf. Jurdy, 1979). Scott et al. (1980), although recognizing this possibility, have constructed arguments which consider the West Philippine Sea as having formed by back-arc spreading but related to the Oki-Daito Ridge as a potential early Tertiary island arc (Karig, 1975) rather than to the Palau-Kyushu Ridge.

Magnetic lineation patterns in the West Philippine Basin have been studied by Loudon (1976), Watts et al. (1977), and Shih (1980) and, together with earlier evidence from deep sea drilling (Sites 290-295; Karig, Ingle, et al., 1975), suggest that active spreading occurred in the early Tertiary (~62-39 Ma), with the NW-SW-trending Central Basin "Fault" representing the extinct spreading center (Ben-Avraham et al., 1972). The younger magnetic anomalies (18-20) are particularly well defined (Watts et al., 1977) and suggest a (half-)spreading rate of 4.4 cm/y.; the older anomalies, though less well defined, indicate a much faster spreading rate. Andrews (1980) has suggested that the spreading direction may have changed 54 Ma from northeasterly to northerly.

Andesite, basalt, and granodiorite have been recovered from the Oki-Daito Ridge, and andesite, diorite, hornblende schists, and serpentine have been dredged from the Daito Ridge; justifiably it has been suggested that these two ridges represent old island arcs (Murauchi

et al., 1978; Karig, 1975; Mizuno et al., 1975, 1979). However, there is no ridge feature south of the Central Basin Ridge ("Fault") which could represent a remnant arc or rifted counterpart of the Oki-Daito Ridge and which would allow for the West Philippine Basin to have formed by back-arc spreading behind the Oki-Daito Ridge during the early Tertiary (Karig, 1975). Scott et al. (1980) have also pointed out that there is no positive evidence from DSDP Holes 294 and 295 that the Oki-Daito Ridge was an active arc during the early Tertiary. Nor indeed do DSDP Sites 445 and 446, close to the Oki-Daito Ridge, record any evidence for arc activity as long ago as the early Eocene (Klein, Kobayashi, et al., 1980). Scott et al. (1980) note that the 6000-meter water depth in the West Philippine Sea is deeper than that for normal ocean crust of early Tertiary age as predicted by the depth-time curve but does appear to fit the depth-time relationship of back-arc basin crust (Sclater et al., 1976). Evidence for the nature of the West Philippine Sea is therefore conflicting.

Site 447, in the eastern part of the basin, lies between the Central Basin Ridge and the Palau-Kyushu Ridge. Paleontological evidence from the basal breccias suggests an age of 32-34 m.y. (Kroenke, Scott, et al., 1980), although the magnetic lineation patterns indicate an age at least 4 m.y. older; there may thus be an appreciable erosion surface between the breccias and the basaltic basement. These breccias are overlain by over 60 meters of middle to upper Oligocene tuffs and polymict breccias representing a volcanoclastic apron derived from the Palau-Kyushu Ridge. The mineralogy, petrology, and geochemistry of Site 447 basalts have been described by Matthey et al. (1980) and Wood, Matthey et al. (1980). They recognize two distinct magma types with different major element glass compositions and with different trace element ratios (e.g., Ti/Zr, Y/Zr) which appear to have evolved in separate magma chambers but nevertheless must have existed simultaneously in order to account for the observed downhole chemical variations. Each magma chamber was periodically replenished with a porphyritic primitive magma and underwent mixing followed by variable degrees of fractional crystallization prior to eruption. However, the overall geochemical characteristics of these basalts—low K, Rb, Ba, Sr, low Nb and Ta, variable light REE depletion, and low  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (~0.7026; Armstrong and Nixon, 1980)—are essentially similar to normal MORB. Abundances of iron as  $\text{FeO}^*$  and  $\text{Al}_2\text{O}_3$  resemble those of MORB rather than back-arc basin basalts, using criteria established by Fryer et al. (this volume). There are thus no obvious arc-related geochemical characteristics in Site 447 basalts such as have been found in some other marginal basins (cf. Tarney et al., 1981; Saunders et al., 1980). There is therefore no *a priori* reason to invoke back-arc spreading for the formation of the West Philippine Sea, although such an origin is by no means excluded.

Basalts recovered from Site 446 in the Daito Basin, between the Daito and Oki-Daito ridges (Fig. 1), are very different from those at Site 447. They occur as a series of sills of post-early Eocene age (Marsh et al., 1980). There are two distinct types of basalt, one rich in



kaersutitic hornblende, the other hornblende-free. Both types have strikingly high contents of titanium (up to 4.9%  $\text{TiO}_2$ ) and incompatible elements as well as relatively strongly fractionated REE patterns ( $\text{Ce}_N/\text{Yb}_N \sim 3-6$ ). Zr/Nb ratios are low ( $<10$ ) but K/Rb ratios high. Many of the basalts are fairly iron-rich, but some are also Mg-rich ( $\sim 15\%$  MgO); although the latter may have high MgO contents as a result of intrasill fractionation, they are still markedly enriched in incompatible trace elements. The basalts are tholeiitic or mildly alkaline. Their geochemical features are not unlike "enriched" MORB from Iceland or from  $45^\circ\text{N}$  on the Mid-Atlantic Ridge (Wood et al., 1979; Tarney et al., 1979) although the latter are not Ti-rich. They are an unusual suite of rocks. Their high Nb and Ta contents (Marsh et al., 1980; Wood, Matthey et al., 1980) contrast with the low Nb and Ta contents of almost all arc lavas and indeed with all marginal basin basalts so far described. Even the mildly alkaline lavas of Penguin volcano in Bransfield Strait marginal basin, next to the Antarctic Peninsula, for instance, have very low Nb contents and high Zr/Nb ratios (Weaver et al., 1979).

It is clear that in their high concentration of Ti, Nb, and Ta these Site 446 lavas have no arc-related geochemical characteristics. The basement age at this site (53 m.y., Klein, Kobayashi et al., 1980) of course predates the development of the Palau-Kyushu Arc. Whereas the petrogenesis of the Site 446 lavas must remain enigmatic, it is evident that there is no basis for linking them in any way to back-arc spreading. It seems more likely, all evidence considered, that the West Philippine Sea is a "trapped" marginal basin unconnected with back-arc spreading in the strict sense.

If, as has been suggested, the Palau-Kyushu Ridge marks the site of an earlier N-S transform which became a convergent plate boundary coincident with the change in direction of Pacific plate motion approximately 45 Ma, the presence of the spreading center with its buoyant hot mantle on the western side of the transform would certainly have facilitated subduction of the cooler Pacific lithosphere on the eastern side of the transform.

### The Palau-Kyushu Ridge

The Palau-Kyushu Ridge is over 2000 km long (Fig. 1) and rises to over 2000 meters above the adjacent basin floors. At Site 448, over 600 meters of volcanic basement consisting of lava flows, dykes, and sills interbedded with volcanoclastic breccias were penetrated below middle Oligocene oozes. Like the dredged samples from this locality (Anonymous, 1977), the basalts are frequently highly vesicular, suggesting eruption at relatively shallow depths or even subaerially. These permeable rocks have been rather readily altered (Matthey et al., 1980), with the clay mineral alteration products being kaolin-rich in the upper part of the core but with mixed-layer smectite-vermiculite-chlorite clays deeper in the section (Aldrich et al., 1980). Nonetheless, the essential geochemical characteristics have been preserved. Nannofossil assemblages and  $^{40}\text{Ar}/^{39}\text{Ar}$  dates indicate that the main period of volcanism began to fade about

32 Ma and had ceased altogether by 29 Ma (Scott et al., 1980). Assuming contemporaneity with the fore-arc lavas on Saipan and Guam, which have yielded ages of approximately 42 m.y. (cf. Ingle, 1975) it would appear that volcanism on the Palau-Kyushu Ridge spanned a period of about 10 m.y. and began within a few million years after the start of subduction.

The geochemical characteristics of the Site 448 lavas has been described by Matthey et al. (1980) and Wood, Matthey et al. (1980). Almost all the lavas are basaltic and conform to the island arc tholeiite series of Jakes and Gill (1970). They are readily distinguished from the basalts of the adjacent marginal basins by their low Cr (20–50 ppm) and Ni (0–20 ppm) contents.

The Palau-Kyushu lavas are quartz-normative tholeiites and basaltic andesites with 48–56%  $\text{SiO}_2$ , 13–17%  $\text{Al}_2\text{O}_3$ , and 10–15% iron as total FeO. FeO/MgO ratios range from 1.5 to over 4.0, much higher than in the basin basalts, implying that they may have undergone considerable crystal fractionation. Currently popular models for the island arc tholeiite series invoke hydrous melting of the mantle wedge above the subducting slab with considerable crystal fractionation during ascent to explain the wide range of Fe/Mg ratios and low Cr and Ni contents. Arc tholeiites from the Palau-Kyushu Ridge have similar geochemical characteristics to those from the Mariana fore-arc (Site 458 and 459) but include more iron-rich compositions. Matthey et al. (1980) noted, however, that the concentrations of many incompatible elements in the Palau-Kyushu lavas, especially the REE, Zr, Nb, and Ta, were much lower than in MORB with equivalent Fe/Mg ratios, and that these did not increase systematically with Fe/Mg ratio. Moreover, they distinguished several petrographic types (ol-plag, opx-cpx-plag, and cpx-plag phyric) that could not be related to each other by fractional crystallization. Thus either several distinct magma chambers were feeding lavas to the Palau-Kyushu Arc or the lavas were closer to "primary" melt compositions than has hitherto been admitted. The only source able to produce such iron-rich compositions would be the subducted ocean crust itself. Although thermal models for steady-state subduction zones (Anderson et al., 1976) tend not to favor melting of the subducted ocean crust, it is to be remembered that the Palau-Kyushu Arc developed rapidly following the start of subduction beneath the West Philippine Sea. Under such conditions high degrees of melting of ocean crust may have been possible. This would account for the consistently high Fe/Mg ratios and low Cr and Ni contents of the arc tholeiites. In many respects the arc tholeiites resemble normal MORB in their trace element characteristics but tend to have rather high LIL element contents, as is typical of most subduction zone magmas.

### West Mariana Ridge

The West Mariana Ridge (Fig. 1) is shallower than the Palau-Kyushu Ridge, usually less than 2000 meters below sea level. Drilling at Site 451 penetrated 930 meters of volcanoclastic material, the breccia fragments being composed mainly of basalts and basaltic andesites with

rare andesites and having phenocrysts of plagioclase  $\pm$  clinopyroxene  $\pm$  orthopyroxene  $\pm$  olivine  $\pm$  magnetite (Mattey et al., 1980; Ishii, 1980).

Chemically these lavas have low contents of Cr, Ni, Zr, Nb, and Ta, like the Palau-Kyushu arc tholeiites, but have low Fe/Mg ratios (1.0–2.0), are lower in both FeO and MgO, have much higher  $\text{Al}_2\text{O}_3$  (18–20%) and CaO (9–13%), and a wider range of  $\text{SiO}_2$  (45–58%). There are larger trace element differences. La/Zr and Ce/Zr ratios are higher and REE patterns are more fractionated, with  $\text{Ce}_N/\text{Yb}_N \sim 2$ . Concentrations of Ba (110–450 ppm) and Sr (450–600 ppm) and other LIL elements are notably higher than in arc tholeiites or in normal MORB. These characteristics are those of calc-alkalic magmas, even though most of the lavas are still basaltic in terms of major elements. Essentially similar, although metamorphosed, rock types were cored at Site 453 during Leg 60 but include cumulus gabbros as well.

### Distribution of Mariana Arc Magma Series in Time and Space

In order to summarize the petrologic history of the Mariana arc, it is important to establish as carefully as possible the various igneous rock series present on each structural component of the arc system, including the islands which we have not yet discussed. Each of the three arc magma series—boninites, arc tholeiites, and calc-alkalic lavas—can be identified on the basis of major oxide chemistry, trace elements, and isotopic geochemistry and mineralogy. Boninite series lavas are particularly distinctive chemically and petrographically, and there is consequently little uncertainty about their identification. The distinction between island arc tholeiites and calc-alkalic lavas is based increasingly on trace elements, although historically the contrasting degree of iron enrichment between the two has served as the basis for classification and been the starting point for much speculation and experimental study. The original definition for calc-alkalic lavas was based on Peacock's (1931) alkali-lime index (the value of  $\text{SiO}_2$  in a differentiation series in which  $\text{total Na}_2\text{O} + \text{K}_2\text{O} = \text{CaO}$ ). But this is not strictly appropriate for most Mariana arc suites containing andesites (they turn out to be calcic rather than calc-alkalic), and has been largely superseded in contemporary petrological thought by the general idea that island arc andesites which show considerable iron enrichment belong to the island arc tholeiite series (e.g., Kuno, 1968; Jakeš and Gill, 1970; Miyashiro, 1974).

Mattey et al. (1980) and Wood et al. (this volume) have coupled this generalized definition with trace element data, noting that calc-alkalic lavas have considerably higher abundances of the mobile hygromagmatophile (HYG) elements than island arc tholeiites. They employ a ternary plot of Th-Ta-Hf/3 as a discriminant. On such a plot (Fig. 10) the calc-alkalic lavas of the Mariana arc system are considerably closer to the Th apex than arc tholeiites, and both suites are considerably depleted in Ta relative to MORB. On the basis of this and other discriminant diagrams, Wood et al. (this volume) propose that the earliest stages of Mariana arc

volcanism are represented by arc tholeiites, whereas younger lavas in the system are calc-alkalic in composition. That is, considering only data for drill sites, the ancestral Mariana arc, which included the present fore-arc region and the Palau-Kyushu Ridge, had arc tholeiite compositions. The younger West Mariana Ridge and the modern arc, each formed after episodes of back-arc basin spreading, have calc-alkalic compositions. Since each of the successive arcs occupied the same structural position with respect to the fore-arc when they were active, Wood et al. (this volume) infer that much the same zone of mantle was tapped in each case. The transition from arc tholeiite to calc-alkalic compositions therefore reflects a fundamental change in mantle source compositions with time—specifically, steady addition of mobile HYG elements to it, derived from the subducting lithosphere.

In terms of iron enrichment, however, this transition was not so straightforward. On Figure 11A, the average compositions for the fore-arc lava types of Holes 458 and 459B (from Table 2) are plotted as well as data fields for samples from Hole 448 on the Palau-Kyushu Ridge and Hole 451 on the West Mariana Ridge (Mattey et al., 1980). A separate, more restricted field for Hole 448 glasses indicates that many of the lavas there have lost iron and gained magnesium as a result of alteration (Scott, 1980), thus suppressing the original iron enrichment and distorting the data field shown on Figure 11A. Similar effects may have occurred in fore-arc Holes 458 and 459B, but our earlier consideration of Figures 8 and 9 suggests that petrological processes (fractionation of titanomagnetite, different source compositions) rather than alteration prevented significant iron enrichment in those basalts. Still, iron enrichment is evident in some of them, and it is probably fair to describe them as arc tholeiites, even though they are not as iron-enriched as the Hole 448 glasses.

Figure 11B compares lavas of the Eocene Alutom and Miocene Umatac formations of Guam (Stark, 1963) and the Eocene Aimeliik formation of Palau (Mason et al., 1956). The data fields for Holes 448 and 451 are again shown. As discussed earlier, some of the Umatac Formation lavas are boninites. The suite from Palau shows so little iron enrichment that it resembles the classical calc-alkalic Cascades suite (Daly, 1933; Smith and Carmichael, 1968). There is somewhat more iron enrichment in the Umatac Formation, but it is not as extreme as in Hole 448 glasses and appears to straddle the field of calc-alkalic Hole 451 lavas. Figure 11C shows data for Eocene lavas of Saipan, which also have minimal iron enrichment.

Data for the present Mariana arc are depicted on Figure 11D. As noted by Dixon and Batiza (1979) and Stern (1979), the extent of iron enrichment is somewhat variable in the different volcanoes of the modern arc, with some islands, such as Sarigan, which have lavas closely resembling the classical calc-alkalic trend and others, such as Agrigan, having somewhat more iron enrichment.

Finally, all these trends can be compared with the compositions of volcanic glasses derived as ash from the

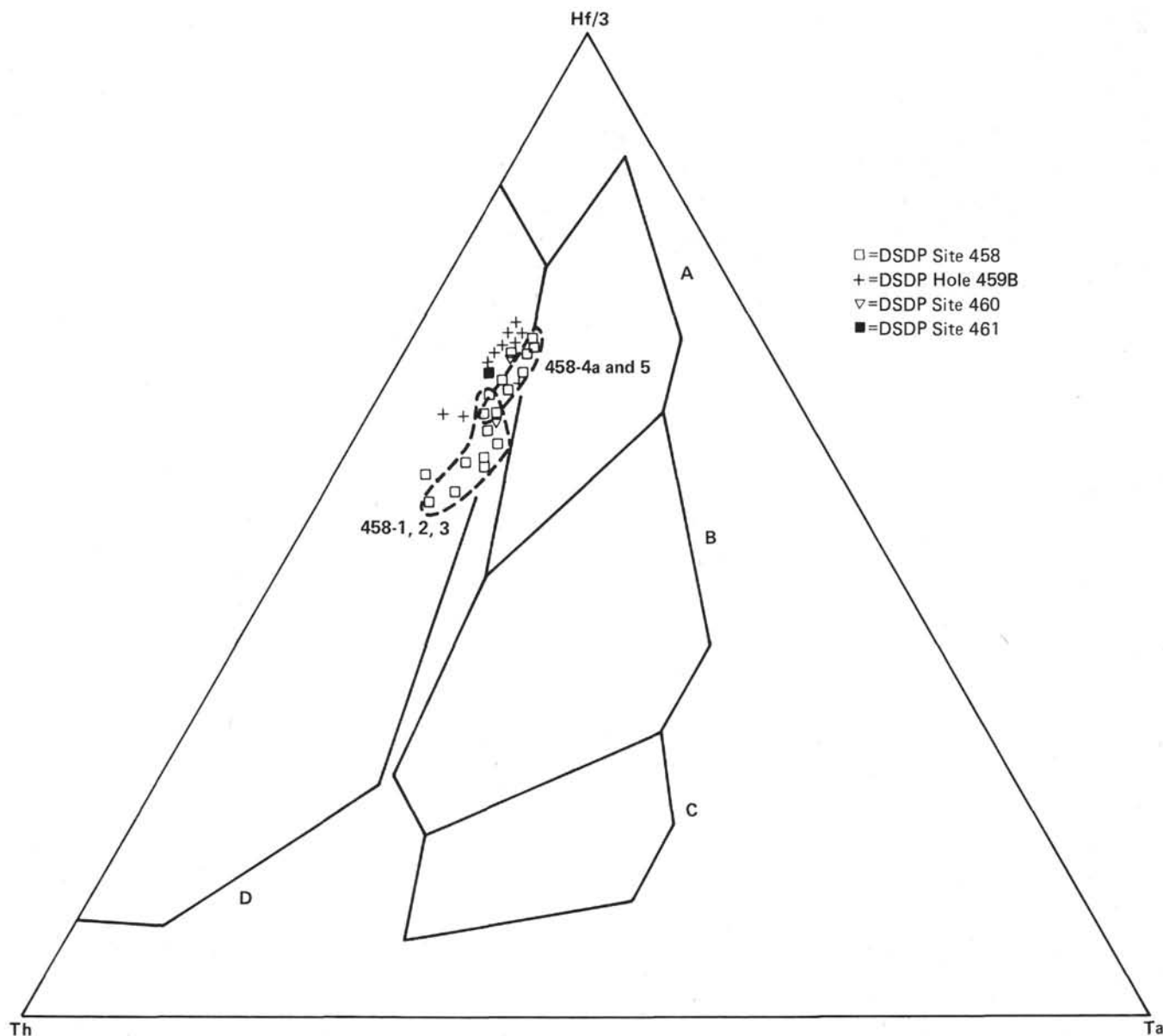


Figure 10. Ternary Ta-Th-Hf/3 diagram for island arc tholeiites and boninites from Holes 458, 459B, 460, and 461 (from Wood et al., this volume). Areas: A = N-type MORB; B = E-type MORB; C = within-plate lavas; D = active plate-margin lavas.

modern arc and deposited as turbidites at Site 453 (Figure 11E, from Packham and Williams, this volume). Many of these glasses show considerable enrichment compared with lavas of the modern arc. Since they also have low  $\text{TiO}_2$ , Packham and Williams describe them as arc tholeiites and note that they erupted from at least 5 Ma to about 200,000 y. ago. The present-day calc-alkalic suites on the modern arc, which show so little iron enrichment, represent a distinct shift in the compositions of lavas erupted on these volcanoes.

The plots of Figure 11 show that eruption of calc-alkalic lavas (those having little or no iron enrichment) and arc tholeiite sequences (with considerable iron enrichment) have varied in time and space in the Mariana arc. This is summarized in Table 4. The ancestral Eocene arc, which included the fore-arc region, Guam, and

Saipan structurally united with Palau and the Palau-Kyushu Ridge, consists of all three major arc rock suites—boninites, arc tholeiites, and calc-alkalic lavas. The Miocene arc, formed after the opening of the Parece-Vela Basin, consists of boninites and calc-alkalic lavas. On Guam these evidently represent a portion of the West Mariana Ridge left behind following opening of the Mariana Trough. The modern arc, which was formed after the opening of the Trough, evidently shifted from arc tholeiite to calc-alkalic compositions in the past 200,000 y.

The question presents itself whether the Eocene calc-alkalic exposures on Guam, Saipan, and Palau have the same trace element characteristics as their younger counterparts on the West Mariana Ridge. Unfortunately, there are no data for the critical trace element discrimi-

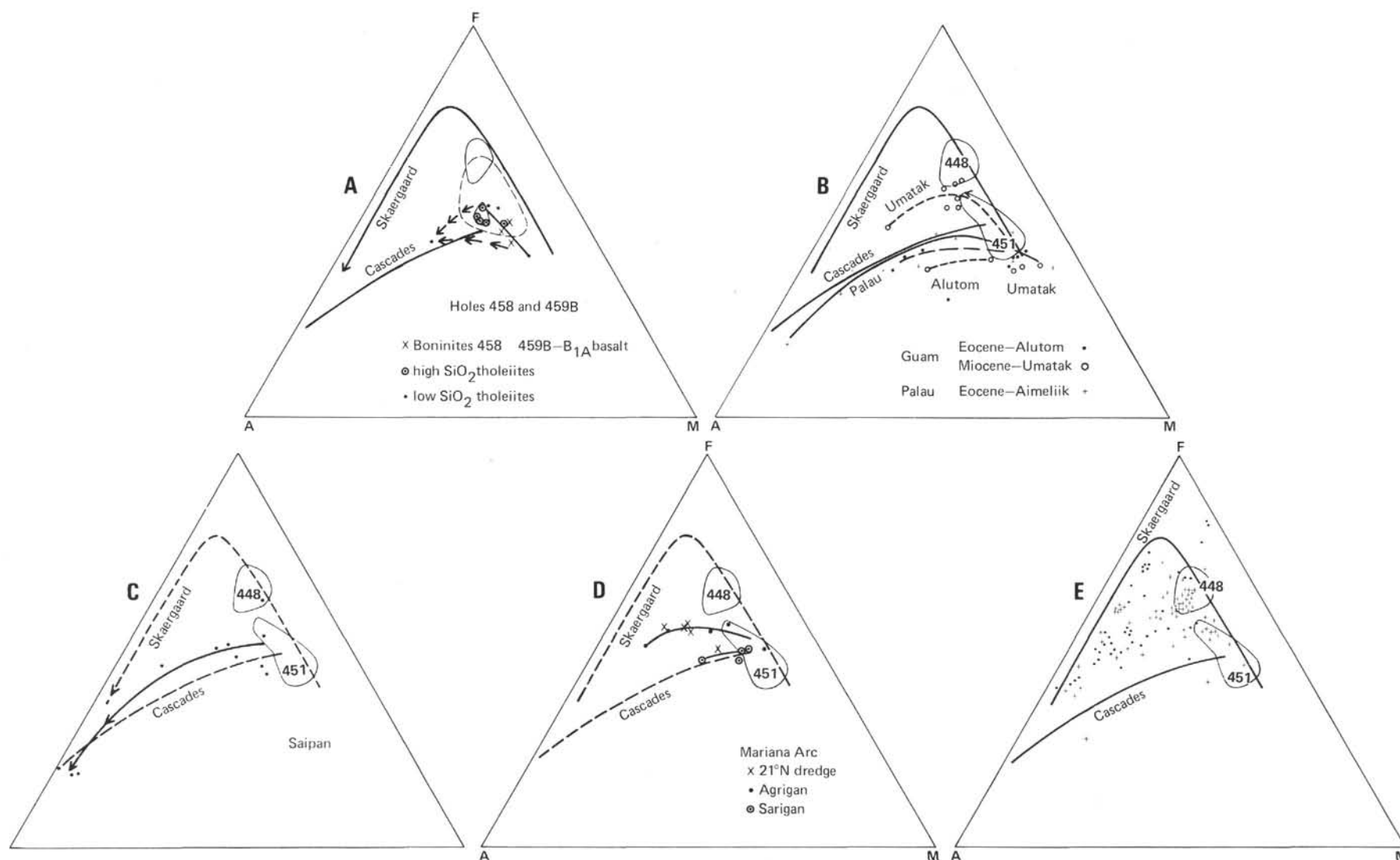


Figure 11. AFM (Na<sub>2</sub>O + K<sub>2</sub>O-FeO\*-MgO) ternary diagrams for tholeiites and calc-alkalic lavas of the Mariana arc system. All diagrams show Skaergaard iron enrichment and Cascades calc-alkalic fractionation trends. A. Holes 458 and 459B, with symbols and groupings as in Figure 8A. (Closed field is Hole 448 glass compositions [Scott, 1980]; dashed field is Hole 448 fresh and altered basalt compositions [Mattey et al., 1980].) B. Eocene and Miocene lavas from Guam (Stark, 1963) and Palau (Mason et al., 1956) compared with fields for Eocene arc tholeiite glasses, Hole 448, Palau-Kyushu Ridge, and Miocene calc-alkalic basalts and andesites, Hole 451, West Mariana Ridge (Mattey et al., 1980). C. Eocene lavas from Saipan (Schmidt, 1957) compared with data fields for Hole 448 glasses and Hole 451 basalts and andesites. D. Modern Mariana arc lavas compared with data fields for Holes 448 and 451 samples. Data for Agrigan are from Stern (1979), for Sarigan, from Dixon and Batiza (1979), and for 21°N dredge haul from Wood et al. (this volume). E. Glass compositions from Hole 453 vitric tuffs derived from the Mariana arc prior to 200,000 y. ago, after Packham and Williams (this volume). Symbols are as in their fig. 5.



Table 4. Known distribution of igneous rock series in the Mariana arc system.

	Sites	Age	Boninites	Arc Tholeiites	Calc-Alkalic Lavas
ancestral Mariana arc	Palau-Kyushu Ridge	448		✓	
	Palau				✓
	Guam				✓
	Saipan				✓
West Mariana Ridge	Fore-arc Trench	458, 459	✓	✓	
		460, 461	✓	✓	
	West Mariana Ridge	451			✓
	Guam		✓		
Mariana Arc		453		✓	
	Sarigan				✓
	Agrigan				✓
	Pagan	457			✓
	Alamagan				✓
	21.4°N				✓

Note: Contrasts between arc tholeiites and calc-alkalic lavas here are based only on the degree of iron enrichment in differentiation series, not to differences in trace element abundances or to parental magma compositions. See text and Figure 11 for discussion.

nants (Ta, Th, Hf, Nb, etc.) in lavas from these islands. There is some suggestion in the major oxide data, however, that they would not be as enriched in these elements as the younger Mariana arc calc-alkalic suites. We mentioned earlier the high alkali-lime indices of Mariana arc lavas in general, which technically classify most of them as calcic rather than calc-alkalic in composition. They are especially high for these older lavas (Table 4) because of their low total alkalis and, especially, their low  $K_2O$ . There are semiquantitative spectrographic trace element data for some Guam lavas (Tracey and Stark, 1963) which show low Zr and consistently lower Ba in Eocene Alutom lavas than in Miocene Umata Formation lavas. Along the presently active arc, there are considerable variations in  $K_2O$  and related trace elements, contributing to differences in the alkali-lime index, even though all the modern arc volcanoes are considered to be calc-alkalic in their general characteristics (Dixon and Batiza, 1979). We therefore conclude that there are variable enrichments in HYG element abundances in Eocene and modern Mariana arc calc-alkalic rock suites, with the younger lavas more enriched than the older. The development of calc-alkalic differentiation sequences does not seem to be entirely dependent on this aspect of their chemistry. The conclusion of Wood et al. (this volume) regarding steady enrichment of mantle sources in the mobile HYG elements may still be valid, although we need the relevant trace element data on samples from Guam, Saipan, and Palau to confirm this.

#### Origin of Calc-Alkalic Magmas in the Mariana Arc System

The origin of andesites and the calc-alkalic magma series remains a major petrogenetic problem. Two aspects of this problem stand out in the context of the Mariana arc: the role of fractional crystallization and the question of whether calc-alkalic lavas represent some sort of evolutionary sequence following island arc tholeiites in time and space.

Andesites and other siliceous lavas are subordinate in volume to basalts in much of the Mariana arc (Larson et al., 1974). Of the islands discussed earlier, only Saipan has a dearth of basalts. Where basalts are abundant,

fractional crystallization tends to become an attractive hypothesis, especially where quantitative modeling supports it. This is the case for the Recent Mariana volcanoes (Stern, 1979; Dixon and Batiza, 1979). Stark (1963) also favored a fractional crystallization mechanism for each of the two sequences (Alutom and Umata) of Guam. Only Schmidt (1957) expressed reservations about the efficacy of fractional crystallization in producing the dacites of Saipan.

Stern (1979) has provided more than the usual range of evidence in support of calc-alkalic fractional crystallization as the origin of the andesites of Agrigan volcano. His study included detailed examination of cumulus gabbros found as xenoliths in the lavas. The xenoliths matched the compositions of a gabbroic mass he calculated to underlie the volcano at depths of less than 5 km. Similar gabbros were recovered in great abundance in association with calc-alkalic metavolcanic rocks derived from the West Mariana Ridge in Hole 453 during Leg 60. In Part I we summarized evidence that the Agrigan xenoliths and Hole 453 gabbros are among a class of gabbros uniquely suited to be the cumulus residua of calc-alkalic fractionation. They have extremely calcic plagioclase, with resulting low whole-rock  $SiO_2$ , and crystallized at high  $P(H_2O)$ , as indicated by their mineralogy. Fractionation of such mineral assemblages in the proportions seen in the gabbros would lead to significant enrichment in  $SiO_2$  with little change in  $FeO^*/MgO$  (the calc-alkalic trend).

From the occurrence of the gabbros in Hole 453, on Agrigan, and even, as argued earlier, south of Guam in the Mariana Trench (Dietrich et al., 1978), it seems evident that they are a major part of the deep crustal structure of the Mariana arc system. Such gabbros have been found in all three of the age-structural domains of the Mariana arc (Eocene, Miocene, and Recent), all of which have calc-alkalic volcanics (Table 4). This suggests that a similar type of fractional crystallization has been a major factor in the production of calc-alkalic andesites throughout the history of the arc.

Two contributors to this volume suggest that calc-alkalic lavas follow extensive arc tholeiite volcanism in the Mariana arc system. Wood et al. view this as a general shift affecting the two younger arc ridges, as we

have discussed, but it seems now that calc-alkalic lavas and their apparently complementary cumulus gabbros occurred much earlier in the history of the arc than suggested by DSDP data alone. Packham and Williams concern themselves only with the presently active arc and propose that the Recent calc-alkalic high volcanoes of the arc have followed arc tholeiite volcanoes which erupted over a much longer time span. Although there are inadequate age data to support it, it is possible that the high volcanoes of each of the older arcs also evolved from an arc tholeiite to a calc-alkalic stage. In the case of the Eocene arc complex, this is at least a spatial, if not a temporal, phenomenon. The subaerial and shallow submarine Eocene lavas of Guam and Saipan clearly erupted on structures higher than the pillow lavas at forearc Sites 458 and 459 and contrast with the latter in having calc-alkalic rather than arc tholeiite compositions. The data for the West Mariana Ridge are the least comprehensive, since calc-alkalic lavas only have been recovered from two closely spaced sites (451 and 453).

In terms of the fractional crystallization mechanism outlined here for producing Mariana arc calc-alkalic andesites (segregation of cumulus assemblages comparable to Hole 453 gabbros and Agrigan xenoliths), a plausible evolutionary sequence can be proposed. The most important condition for such fractionation to occur is development of magma chambers in the deep crust at high water pressures. Gravity and seismic data demonstrate that the crust beneath the fore-arc region near Sites 458 and 459 is thin (LaTraille and Hussong, 1980; Hussong and Fryer, this volume). This is verified by the persistent recovery of ultramafic rocks at fairly shallow levels at many places in the Mariana Trench (Dietrich et al., 1978; S. Bloomer and J. Hawkins, personal communication). Fore-arc magma chambers thus were shallow, and the conditions of high  $p(\text{H}_2\text{O})$  required for calc-alkalic fractionation could not be achieved. Beneath lines of subaerial arc volcanoes, however, the crust is considerably thicker (see Plate 2, back pocket, this volume; Hussong, this volume). The crust beneath each of the arcs must also have thickened with time. The coarse and highly porous pyroclastic ejecta of even basaltic island arc volcanoes guaranteed that considerable amounts of water were trapped within thickening lava piles. Even if the same magma chambers continued to exist, they had to become more deeply buried as the volcanoes built. There should thus have been a tendency for tholeiitic island arcs to develop conditions suitable to calc-alkalic magmatism. This need not have happened simultaneously along the length of the arcs, but ultimately it should have happened generally along all of them.

### Back-arc Basin Opening and Arc Volcanism

Recovery of gabbros at Site 453 demonstrated that opening of the Mariana Trough split the West Mariana Ridge deep into its crustal roots. Solidified magma chambers were torn in two. There is no question that this must have profoundly affected arc volcanism. Scott et al. (1980) and Scott and Kroenke (1980) argue, on the basis of available radiometric and paleontologic data

and on the occurrence of ash beds in sediments cored in the Parece Vela Basin, that arc volcanism waned considerably, or ceased altogether, with each opening of a back-arc basin in the Philippine Sea. Clearly, however, following opening of the Mariana Trough, arc volcanism soon resumed. Vitric turbidites derived from air-fall volcanic ash produced at the presently active Mariana arc have been deposited at Site 453 nearly continuously for the past 5 m.y. (Packham and Williams, this volume). Upper Miocene-lower Pliocene hiatuses in fore-arc Holes 458 and 459B (Kling, this volume; Ellis, this volume) suggest that the earlier stages of opening of the Mariana Trough were unaccompanied by arc volcanism. However, there is no corresponding upper Oligocene-lower Miocene hiatus in Hole 459B for the opening of the Parece Vela Basin. Some ash was deposited in every core in this time interval, but since the sediments of Hole 459B are primarily turbidites, it is possible that erosion and sediment transport processes simply persisted throughout any interruption in volcanism on the arc.

Rodolfo and Warner (1980) suggest that the earliest stages of back-arc rifting consist of mixed basalt-andesite volcanism because emplacement of basaltic magmas in contact with water-saturated pyroclastic debris shed into the newly opened narrow rift. Once this happened, however, what were the compositions of the earliest lavas to erupt along the new arcs? Packham and Williams (this volume) think that the compositions of the Mariana arc and Mariana Trough lavas have diverged in the past 5.0 m.y. from a single basaltic composition resembling Trough basalts. This makes sense in that the magma sources for the present Trough and arc are still quite close together (approx. 120 km) and probably were only a few tens of kilometers apart when the line of the arc first became well established. Wood et al. (this volume) and Fryer et al. (this volume), however, note that basalts in the Trough have some of the geochemical characteristics of arc lavas, notably enrichments in some of the mobile HYG elements. In the earliest stages of rifting, such enrichments might have been caused either by contamination (e.g., Rodolfo and Warner, 1980) or by HYG-element enrichments in the mantle source. Probably, therefore, the initial arc and Trough lavas were similar but had compositions in between what we recognize as modern arc and Trough basalts. The early proximity of arc and Trough magma sources almost certainly means that arc lavas were then even more dominated by basalts than now. This is borne out for the West Mariana Ridge by an increase in the abundance of andesitic ash and a decrease in hyaloclastic basalt at Parece Vela Basin Site 450 as the basin widened in the Miocene (Rodolfo and Warner, 1980).

From these considerations, it would seem that for calc-alkalic magmatism to have occurred on either the West Mariana Ridge or the Mariana arc, it must have followed greater or lesser outpourings of basalts. With the Trough opening so close to the arc, geothermal gradients beneath the arc, too, would have been steeper and the depth of origin of arc basalts correspondingly shallower. Conditions were favorable to the development of

shallow magma chambers in which iron enrichment fractionation would have occurred. This would have been essentially arc tholeiite magmatism. As the thermal regimes responsible for volcanism along the arc and in the center of the Trough separated, the sources of arc magmas should have deepened, eruption rates diminished, the crust of the arc thickened, and the conditions for calc-alkalic gabbro fractionation approached. A similar sequence should also have occurred on the West Mariana Ridge, following opening of the Parece Vela Basin.

### Formation of Back-arc Basins

Plate tectonic theory has had considerable success in accounting for the principal physiographic features of the ocean floor and the ages of different portions of the ocean crust. One of the great tasks remaining before a comprehensive global tectonic synthesis is achieved is to account for the tremendous quantity and variety of igneous activity and lava types at both accretionary and convergent plate boundaries. Back-arc basins are a particularly complex problem because they form in the midst of the principal physical manifestations of volcanism at convergent plate boundaries, the island arcs. Since we are far from understanding the laws which govern the placement and perpetuation of spreading centers in the ocean basins, these *interarc* basins would seemingly be even more difficult to explain.

The principal hypotheses to date that account for back-arc basins have their considerable parallels in the original hypotheses of seafloor spreading and plate tectonics—viz., the two-limbed convection model of Hess (1962) and, alternatively, the recognition that spreading centers are not directly linked to such a simple system but respond more or less passively either to larger patterns of convection in the mantle or perhaps to the motions of the larger plates (e.g., McKenzie, 1967, 1969). Thus Karig's (1971b) original back-arc basin hypothesis was that a "diapir" of mantle material rose from the subducting lithospheric plate to sunder the arcs and initiate spreading basins by a type of shallow two-limbed convection. Karig supposed that frictional heating at the subduction zone (Oxburgh and Turcotte, 1970) caused the diapir, but more recent models propose simply that subduction drags mantle material downward, forcing convective turnover behind the arc and producing tension and back-arc rifting (Sleep and Toksöz, 1971; Andrews and Sleep, 1973; Toksöz and Bird, 1977; Toksöz and Hsui, 1978). Bibee et al. (1980) have extended Karig's hypothesis to explain apparently asymmetric spreading and ridge crest jumps in the Mariana Trough. These keep the center of spreading close to the Mariana arc above the subduction-triggered diapir. Uyeda and Kanamori (1979), however, argue that western Pacific back-arc spreading occurs in response to retreat of the Eurasian plate from the Pacific plate, which establishes a dominantly tensional stress regime in such regions as the Philippine Sea, allowing back-arc basins to form without the formation of a mantle diapir. There are similar proposals by Chase (1978) and Jurdy (1979).

Any one of these hypotheses can explain how the Mariana arc has twice been split by back-arc basins. The diapiric model does not allow the locus of back-arc spreading to shift too far to the west of the arc and away from the subduction zone. Episodically, a major ridge-crest jump to a position within the arc itself must occur for the zone of rifting to remain centered on the mantle diapir. Toksöz and Hsui (1978) argue that back-arc spreading actually forces the arc against the subducting plate, decreasing its dip and ultimately making overturn of the mantle caused by subduction impossible. Back-arc spreading ceases, the steep dip of the subduction zone is gradually restored, and a new cycle begins. According to the passive spreading model, periods of retreat of the Eurasian plate from the Pacific plate may alternate with periods of relative convergence in which back-arc spreading ceases, so that the two halves of back-arc lithosphere combine into a single "welded" plate. Renewed tensional stress will initiate a new rift along the weakest line of the combined plate. In the absence of an active back-arc spreading center, this is believed by Uyeda (personal communication) to be the line of the arc.

Petrological data presently do little to constrain these hypotheses inasmuch as our understanding of the origin of island arc lavas is sufficiently ambiguous to allow virtually any combination of petrological and tectonic speculation. The subduction zone drag model probably has the greatest difficulties from a petrological viewpoint since it places arc volcanism directly above the downwelling limb of the proposed convection cell, a difficulty recognized by Andrews and Sleep (1973) and Toksöz and Hsui (1978). The direct opposite is implied by petrological models requiring components from the subduction zone either to infiltrate or rise diapirically into the overlying mantle to become sources for arc lavas. Subduction zone drag offers no explanation for the source of heat to produce these lavas. Possibly, no specific addition of heat, produced for example by friction, is necessary. Dick (1980) argues that the high vesicularity of arc and back-arc basalts reflects addition of water to the mantle above the subducting slab. This lowers the solidus, produces melting (e.g., Wyllie, 1971; McBirney, 1965), and reduces mantle density (O'Hara, 1975), causing diapiric upwelling. Melting continues in the rising diapirs until the water-saturated solidus is reached at low pressures.

We have argued here generally in favor of segregation of island-arc-tholeiite, calc-alkalic, and olivine-boninite parental magmas from the relatively shallow mantle beneath island arcs—i.e., not from the subducted slab—in a manner consistent with Dick's hypothesis. There appears to be no way to distinguish on geochemical grounds whether the immediate mantle sources of the various magmas were originally directly beneath the arc or arose by some diapiric mechanism from the slab to the shallow levels where melt segregation took place. At most, we can infer only that the mobile HYG elements were derived from the subducting materials and infiltrated the shallow mantle beneath the arc, causing the



particular enrichments in these elements observed to a greater or lesser degree in all the arc lavas. But whether or not the peridotite component of the magma sources was originally in place beneath the arcs, or rose from the depths, extraction of magmas and the development of a thick island arc lava pile would entail considerable accumulation beneath it of refractory peridotite depleted in basalt or basaltic-andesite fractions parental to arc lavas. Boninites themselves are evidence that just such depleted peridotite exists in reasonable abundance beneath the Mariana fore-arc region and the island of Guam.

We suggest that after sufficient accumulation of such shallow refractory peridotite beneath an island arc, it would become increasingly difficult to extract a melt fraction from it. With the building of high volcanoes, it would also become hydraulically more difficult for that melt fraction to reach the surface. Only the more volatile and explosive differentiated lavas would erupt. We can therefore surmise that conditions in the mantle itself would add to those previously described and favor a shift toward calc-alkalic compositions during the history of an island arc.

We now carry this one step further, to ask whether accumulation of such a refractory mass could also lead to the development of conditions favorable to the initiation of a new back-arc basin. Either some source of heat exists, related to subduction, that produces arc magmas, or, as Dick suggests, the mantle solidus is suppressed by addition of water, resulting in diapirism and melting. In either case, the effects should have become more or less continuous and involved the same zone of mantle to build three successive island arcs in the Marianas system. Volcanism itself represents release of energy, both thermal and gravitational (buoyant). One can readily suppose that as the mantle beneath each arc became steadily more refractory and less susceptible to having energy released from it by volcanism, the energy would manifest itself in other forms.

We suggest, then, that there may have been a type of petrological feedback mechanism in the Mariana arc system related to the long-continued build-up of arc magmas in restricted geographical areas. In appropriate conditions of tensional stress in the region, this feedback was manifested by major reorganizations of the mantle beneath the arcs, wherein the thick, refractory mantle that formed beneath them was split open and replaced by new depleted mantle. This became the source for both back-arc basin basalts and lavas of the new island arc (Fig. 12A). We see this simply as a consequence of buildup of heat and/or buoyant forces beneath thickening zones of mantle that were an increasing obstacle to release of either form of energy to the earth's surface by means of volcanism. With such a mechanism we would expect separate cycles of mantle enrichment in the mobile HYG elements corresponding to the building of each separate arc (Fig. 12B). For reasons given in this and earlier sections, these cycles would be superimposed on the evolution of each arc from arc tholeiite to calc-alkalic compositions.

With passive back-arc spreading models such as that of Uyeda and Kanamori (1979), there would be no

necessary connection between the petrological state of development of an arc (or the mantle beneath it) and the cessation of back-arc spreading in one basin and its renewal in another. Any apparent arc petrologic sequence would be fortuitous. Conversely, lack of a consistent sequence might support such a model. There is potential difficulty, however, in providing a mechanism by which successive arcs are split precisely down their lengths in response to the motions of distant plates. Arc rifting involves stretching, thinning, and ultimately separating the arc crust along one of its thickest parts, tearing into deep, solidified crustal magma chambers. It is not obvious that this should be the weakest part of the crust, even granting a "welded" back-arc lithosphere. The mechanism proposed here does not depend so much on a "weak" zone of lithosphere as on a concentration of forces beneath the arcs to produce rifting. We believe that the calc-alkalic gabbros and metamorphic rocks of Hole 453 in the Mariana Trough, which crystallized and were metamorphosed at fairly substantial depths (5–10 km or more), indicating uplift, imply the action of buoyant forces during this mantle reorganization at the outset of rifting.

Uyeda and Kanamori (1979) attempt to account for voluminous calc-alkalic lavas in eastern Pacific arcs such as the Andes by proposing that a regional compressive stress pattern keeps back-arc basins from developing, preventing eruption of magmas derived at shallow levels in the mantle (i.e., basalts). If this is indeed the case, then one could as well say that the compressive regional stress regimes simply prevent basins from opening, whether or not there are diapirs. The arcs would be forced to stay near or above any zones of concentration of heat or buoyant forces. Since they would have to feed upon themselves for release of energy, voluminous calc-alkalic magmatism might be expected. A tensional regional stress pattern in the western Pacific may therefore contribute to the formation of back-arc basins but does not necessarily preclude convective disturbances in the mantle such as the one depicted in Figure 12. We point out that once the type of convective disturbance envisaged has occurred, new and fertile mantle sources are available for generation of both arc and back-arc basalts for some time to come, without any additional major overturn or counterflow of the mantle.

Future petrologic contributions to hypotheses of Mariana arc evolution and back-arc basin origin must focus on evaluation of the mantle sources of arc lavas, particularly those of the readily available calc-alkalic suites of the older islands and of the earliest stages of volcanism on the presently active arc. The latter may be possible only with deep drilling on one of the modern volcanoes.

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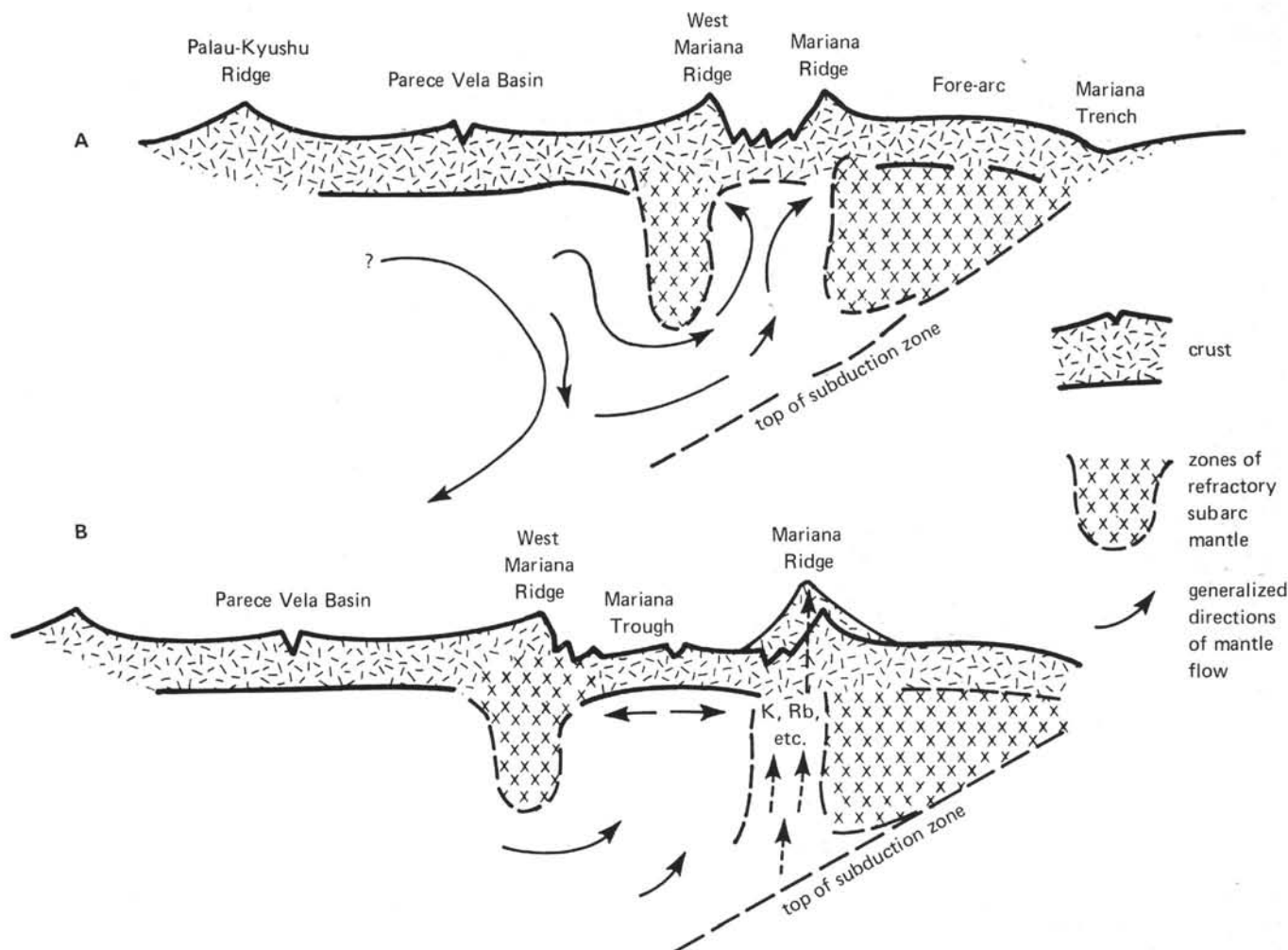


Figure 12. Suggested model for back-arc basin rifting using the Mariana Trough as an example. A. Disruption of refractory subarc mantle and introduction of new fertile mantle in its place. B. Continued back-arc spreading coupled with simultaneous modification of subarc mantle by (1) extraction of magmas at top and (2) addition of mobile HYG elements from below. Eventually (1) outstrips (2), resulting in refractory, infertile subarc mantle from which melts are difficult to extract, necessitating return to A.

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